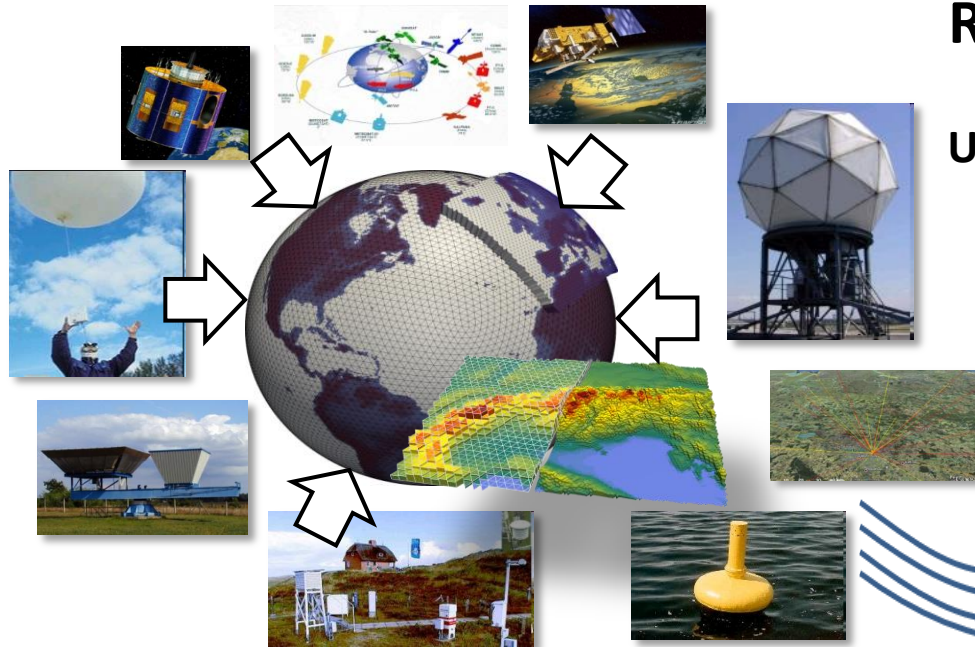


Ensemble Data Assimilation and Particle Filters for NWP

With the help of many people, in particular:

Anne Walter,
Andreas Rhodin
Harald Anlauf,
Christina Köpken,
Robin Faulwetter,
Olaf Stiller,
Alexander Cress,
Martin Lange,
Stefanie Hollborn,
E. Bauernschubert,
Christoph Schraff,
Hendrik Reich,
Klaus Stephan
Ulrich Blahak



Roland Potthast
NWP @ DWD &
University of Reading

Partners include:
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S. Dance
Nancy Nichols

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1. Why and Where **Distributions, Risk** and **Uncertainty?**
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Why Distributions , Risk, Uncertainty?

Extreme Weather Events
triggered by climate change
threaten infrastructure and
economy.

Lightning, Cyclones, Heavy
precipitation events impact
the life of individuals and the
society.

Protect Lives!

Protect Property!

Protect Society and Business

**We need Uncertainty
Estimates for Risk Prediction!**

**National Task: Warn
and Protect**



Phenomena	Level
Small (5-10)	Moderate
Medium (10-25mm/h)	Strong
Storm Force Gusts, Heavy Rainfall	Strong
Storm Force Gusts, Heavy Rainfall, Hail	Strong

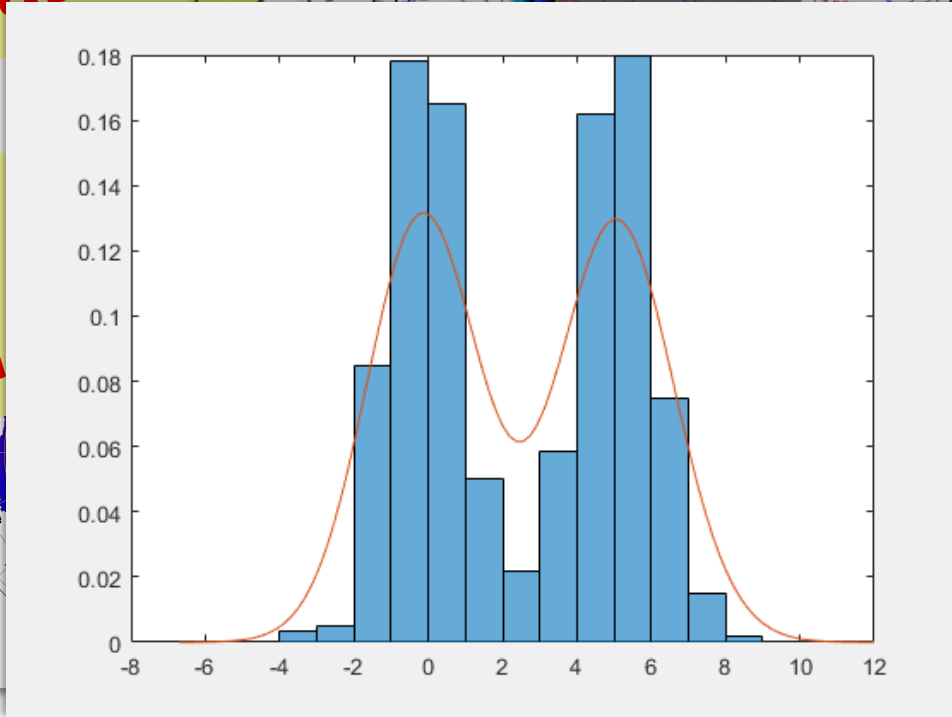
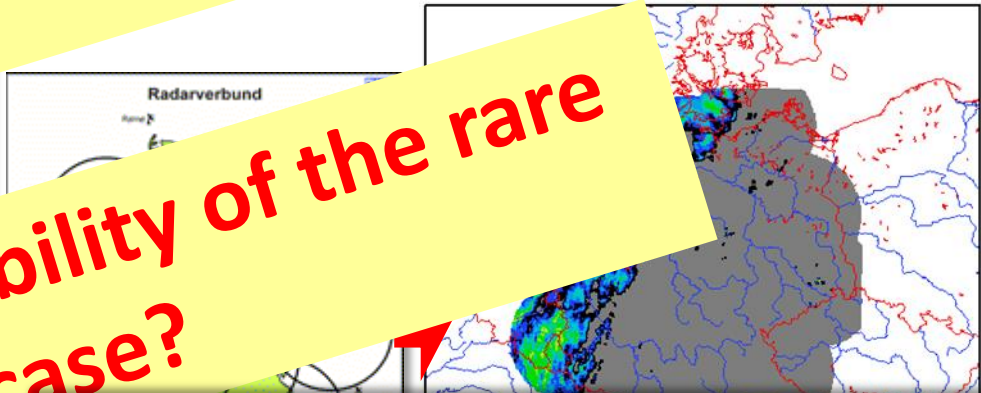
**NOW, 5min, 30min,
... 1h, 2h, 6h, 24h, 72h, ...**

Why Distributions, Risk, Uncertainty?

More than what would usually happen!

What is the probability of the rare event, the worst case?

We need Uncertainty Estimates for Special



Why Distributions , Risk, Uncertainty?

6h Market: Millions

Day-Ahead Market: Billions!

We need Uncertainty Estimates for Energy Network Security!

Renawable Energy Forecasting

Netzebene 1 Übertragungsnetze Höchstspannung 380 kV

Netzebene 3 Überregionale Verteilnetze Hochspannung 110 kV

Netzebene 4 Transformierung

Netzebene 5 regionale Verteilnetze

Sources of UK Electricity 2016

Source	Percentage
Nuclear	27%
Wind	12%
Coal	9%
Biomass	8%
Solar	4%
Hydro	1%

und Export

Höchstleistung-Windparks

Erdgas

Wasserkraft

Überregionaler Ausgleich

Solarpark

Windpark

Industrie

Wärmepumpen

Solaranlagen

BHKW

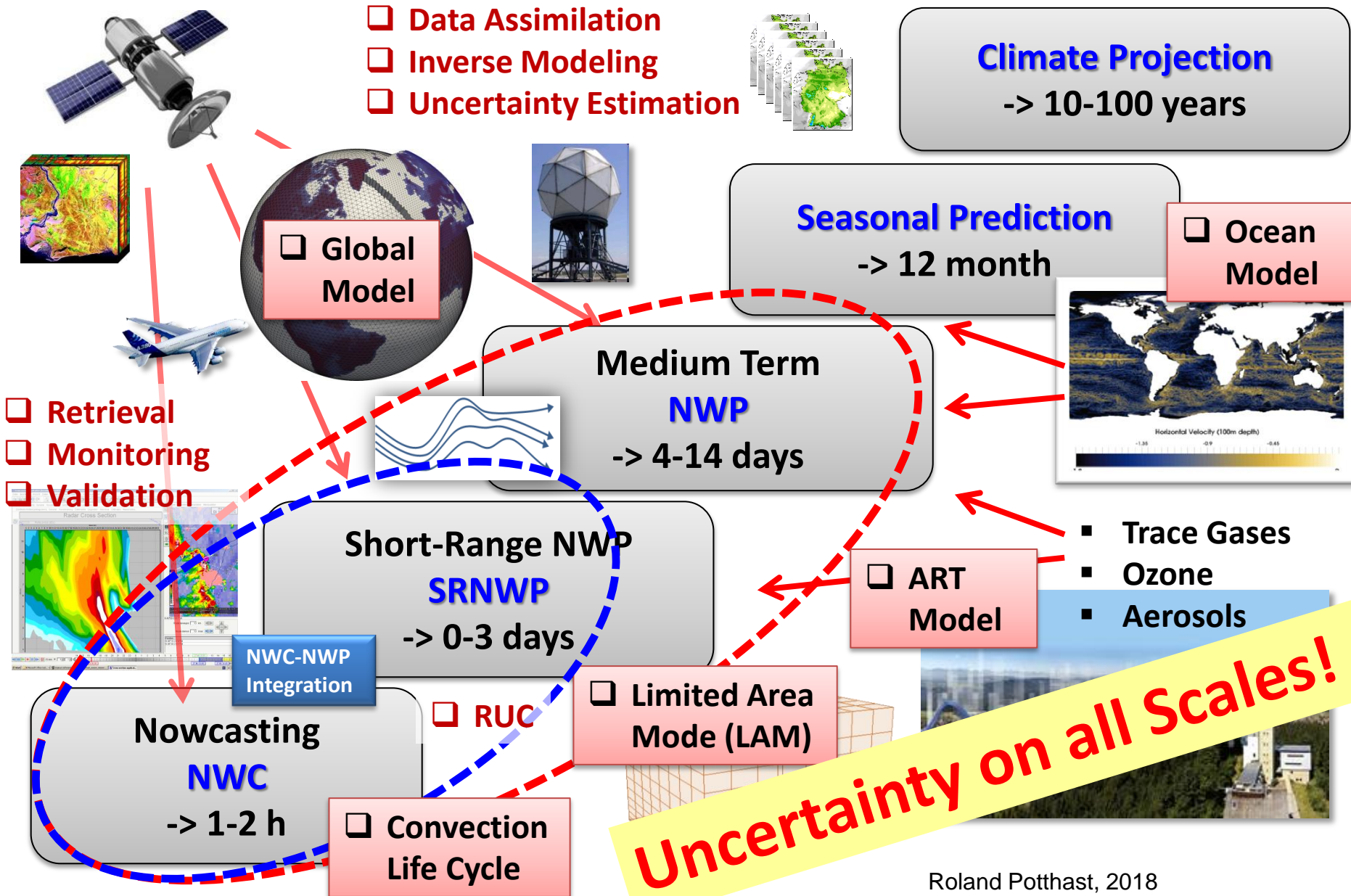
Netzebenen der Stadtwerke

net/pressemitteilung-76

TE

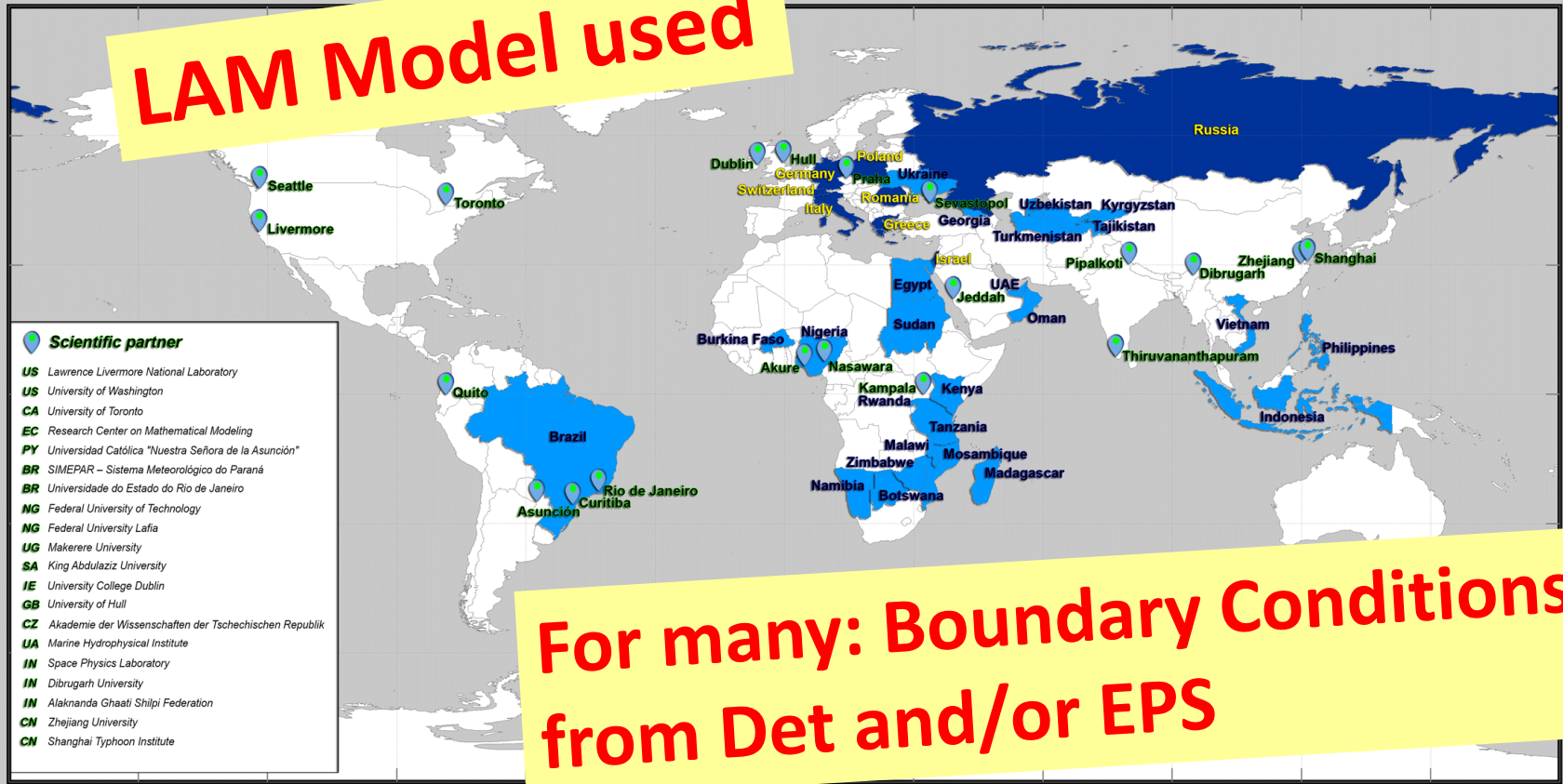
TRANSNET BW

Framework Numerical Weather Prediction



40 Countries

LAM Model used

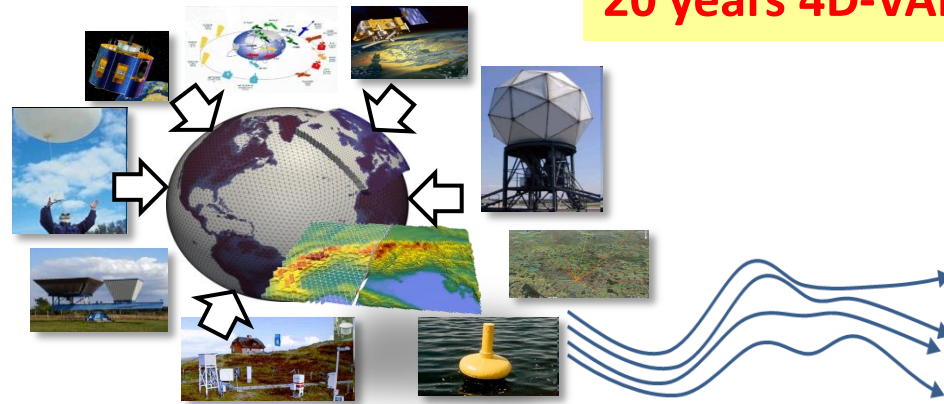


For many: Boundary Conditions from Det and/or EPS

Cosmo User Seminar + Training Course + Symposium

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- Why **variational** Data Assimilation (3D/4D-VAR)?



20 years 4D-VAR at ECMWF

Since 1990

Since 2000

Since 2010

Since 2020

- Why **Ensemble** Data Assimilation (EDA)^(c)?

- Why **Hybrid** Methods? (3D/4D-EnVAR)

- Why Particle **Filters**? (PF,GPF,ETPF,LAPF,LMCPF)

Variational Analysis (3D/4D-VAR)

Recall where we came from ...

The minimization of

$$J(x) = \|x - x^b\|_{B^{-1}}^2 + \|y - Hx\|_{R^{-1}}^2$$

Obs Operator



Climatological covariances

leads to the analysis

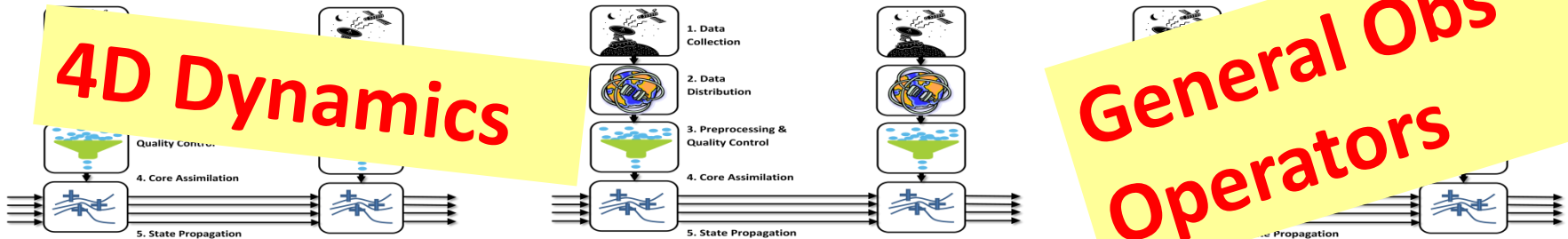
Minimization

$$x^a = x^b + B H^T (R + H B H^T)^{-1} (y - H x^b)$$



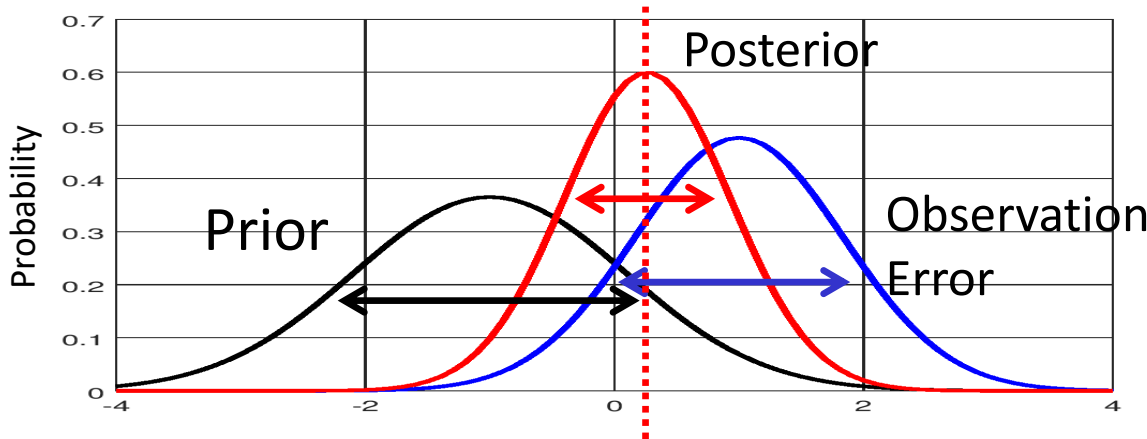
This is the mother of all data assimilation formulas

4D Dynamics



General Obs Operators

Stochastic View \Leftrightarrow Minimization



**Var == Mean/
ML estimator!**

Gaussian Prior

$$p(x) = e^{-\frac{1}{2}(x-x^b)^T B^{-1}(x-x^b)}$$

Gaussian Data Error

$$p(y|x) = e^{-\frac{1}{2}(y-Hx)^T R^{-1}(y-Hx)}$$

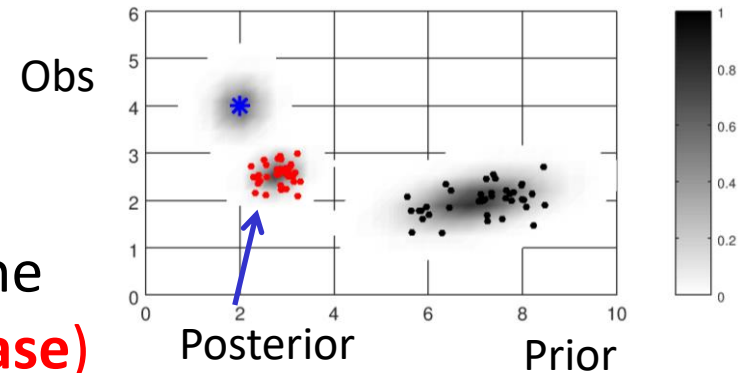
Gaussian Posterior

**Maximum Likelyhood Estimator =
Minimization of Functional = 3/4DVar**

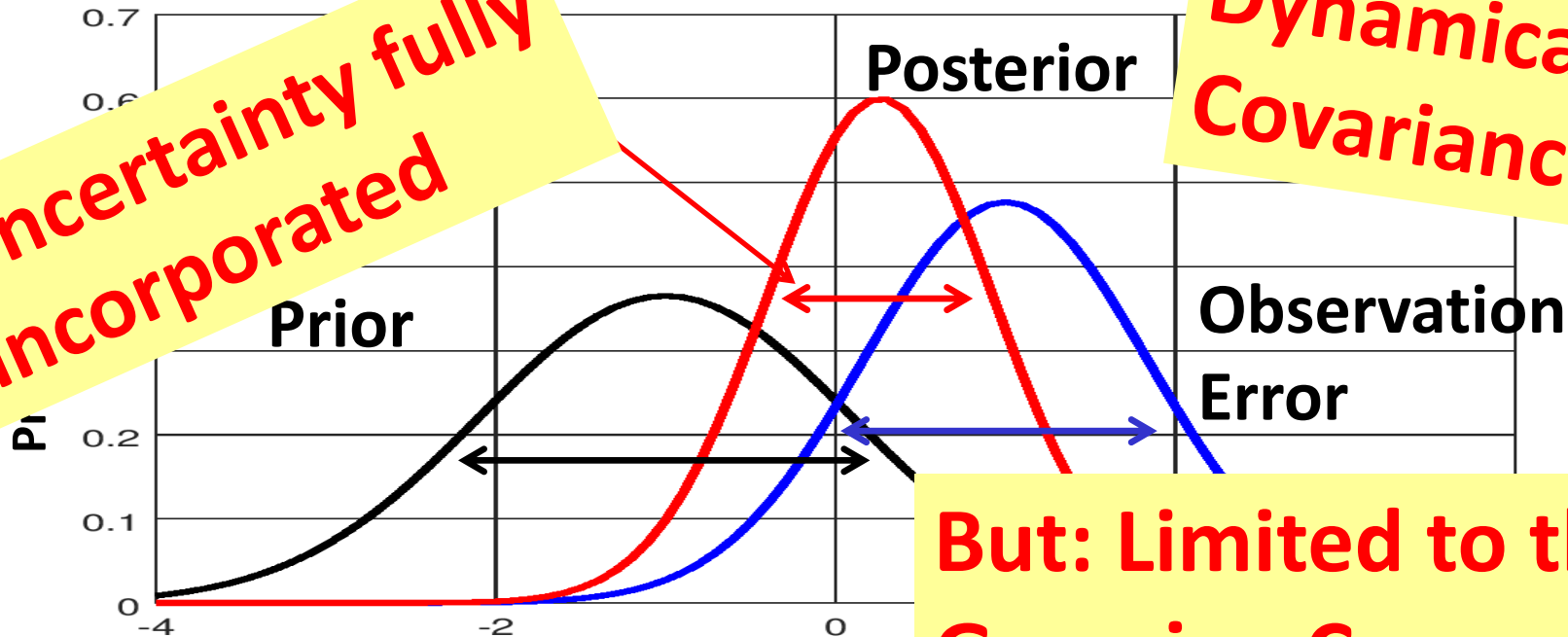
$$p(x|y) = c e^{-\frac{1}{2} \left\{ (x-x^b)^T B^{-1}(x-x^b) + (y-Hx)^T R^{-1}(y-Hx) \right\}}$$

Basic Idea of the Kalman Filter

- **Sequential** Assimilation of Data
- Do not only adapt the mean, but also the **Covariance B (Uncertainty, Gaussian Case)**



Uncertainty fully incorporated

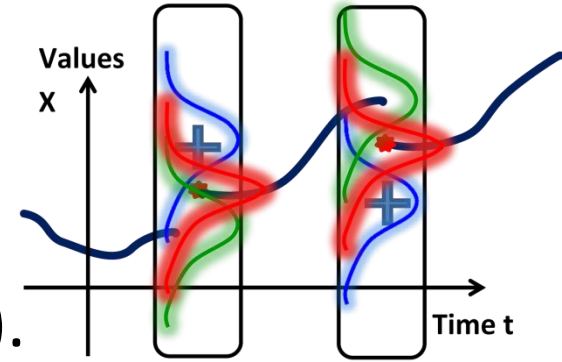


Dynamical Covariances

But: Limited to the Gaussian Case

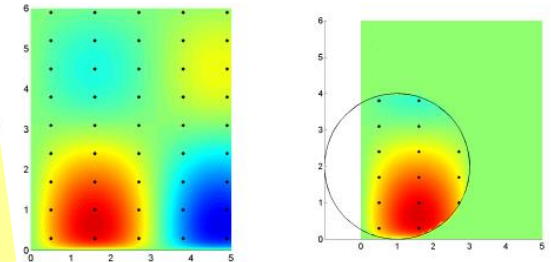
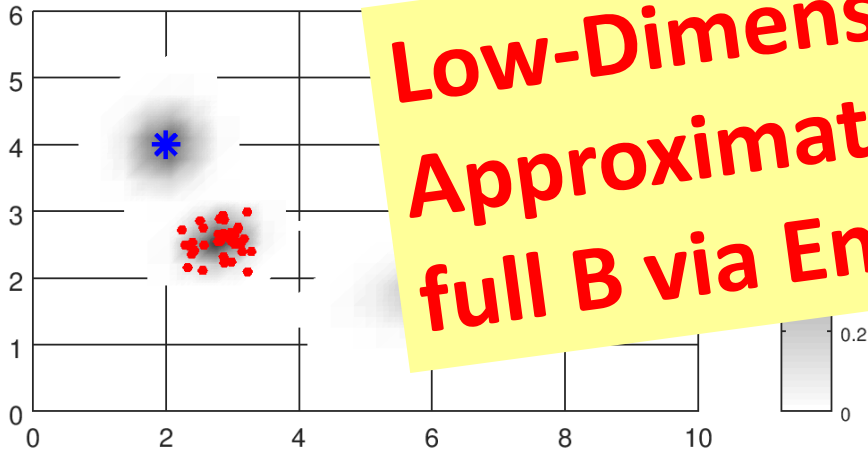
EDA: Ensemble Kalman Filter (EnKF)

- Kalman Filter needs B update => **expensive!**
- **Estimate** B based on an ensemble of forecasted states (**stochastic estimator**).



B will be **flow-dependent** and variable, depending on the **model dynamics** and on the **observations**

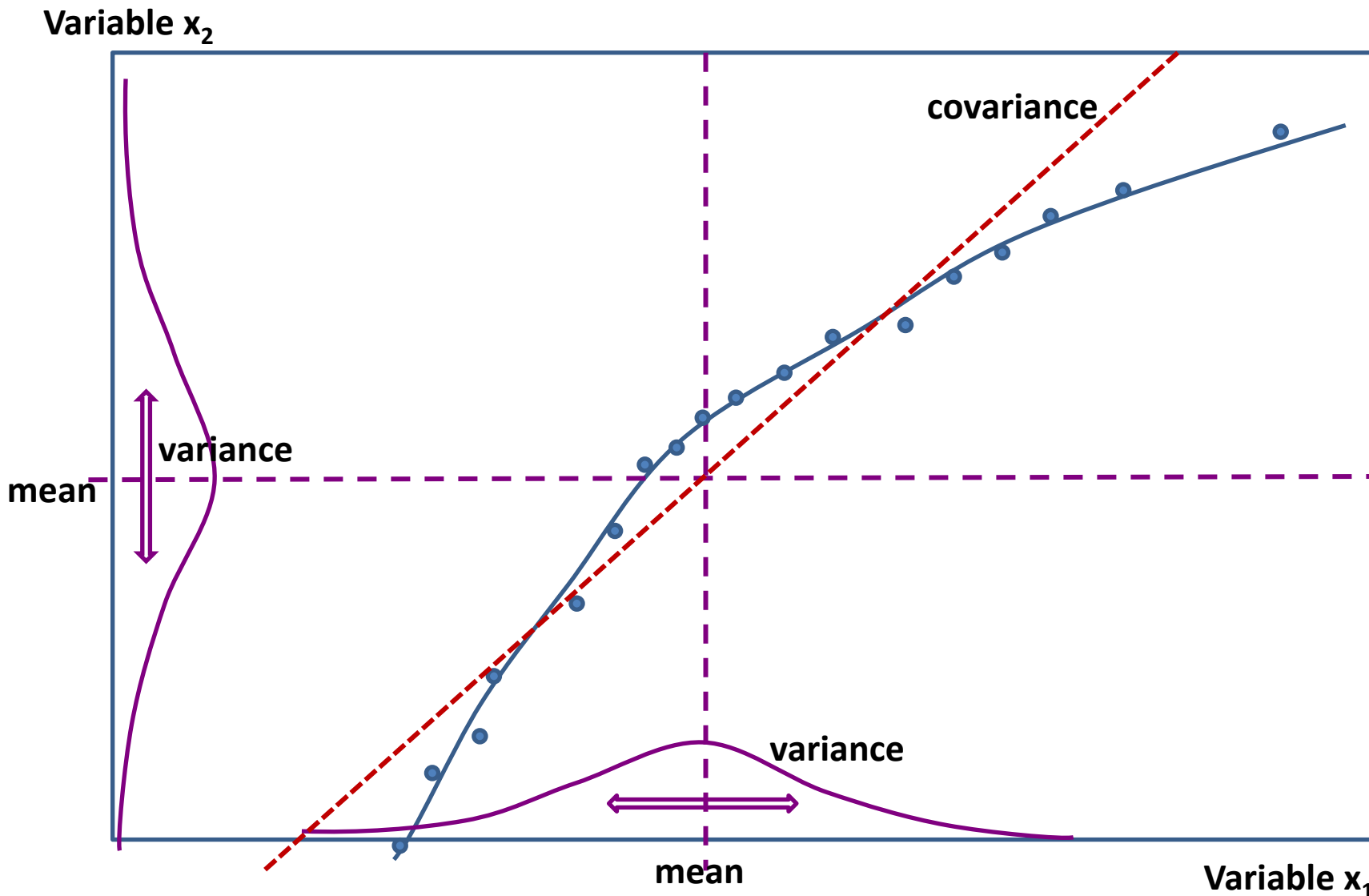
**Low-Dimensional
Approximation to
full B via Ensemble!**



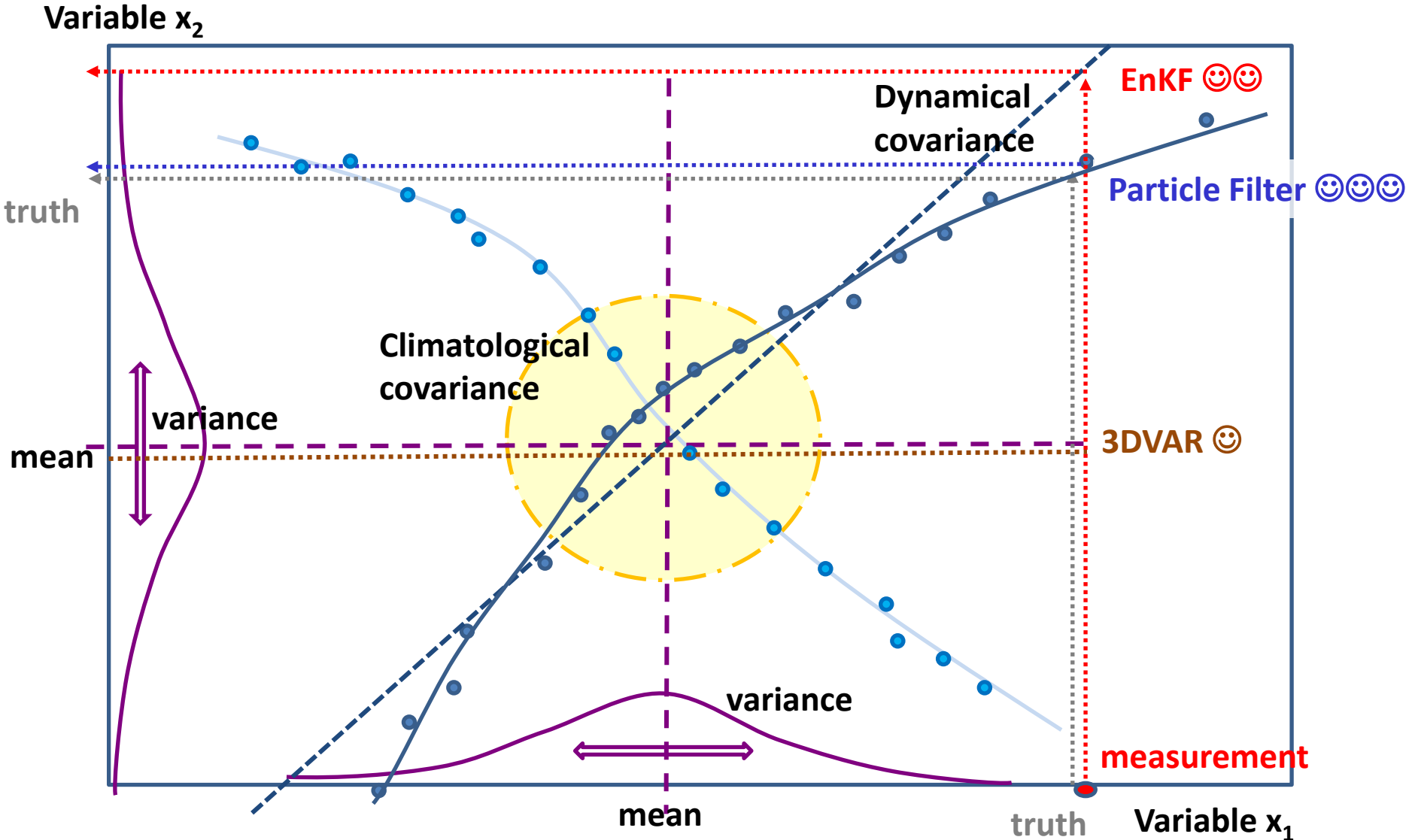
Needs Localization

Localized: LETKF

Using Perturbations, Climatological or Dynamical Covariances, and the Particle Filter



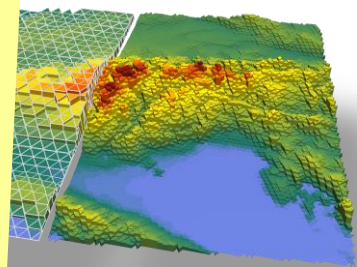
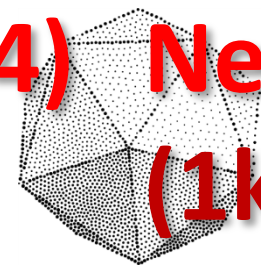
Using Perturbations, Climatological or Dynamical Covariances, and the Particle Filter



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- 1) ICON Model 13km
- 2) Nest over Europe (6.5km; 2-way)
- 3) ICON-LAM D2
- 4) Nest over Germany (1km; 2-way) D1
- 5) NWC Ensemble



- 3h cycles global+Europe: 180h fc
- 1h ana cycle LAM, 3h fc cycle: 27h fc
- RUC cycle 1h: 6-12h fc

$$\frac{\partial v_n}{\partial t}$$
$$\frac{\partial w}{\partial t}$$
$$\frac{\partial \rho}{\partial t}$$
$$\frac{\partial \rho v}{\partial t} + \nabla \cdot (\bar{v} \rho \theta_v) = 0$$

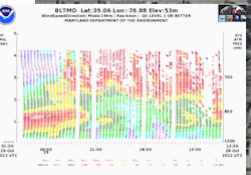
Conventional and Remote Sensing

Exploring Pilot Stations for Boundary Layer Remote Sensing Obs 2019-2025

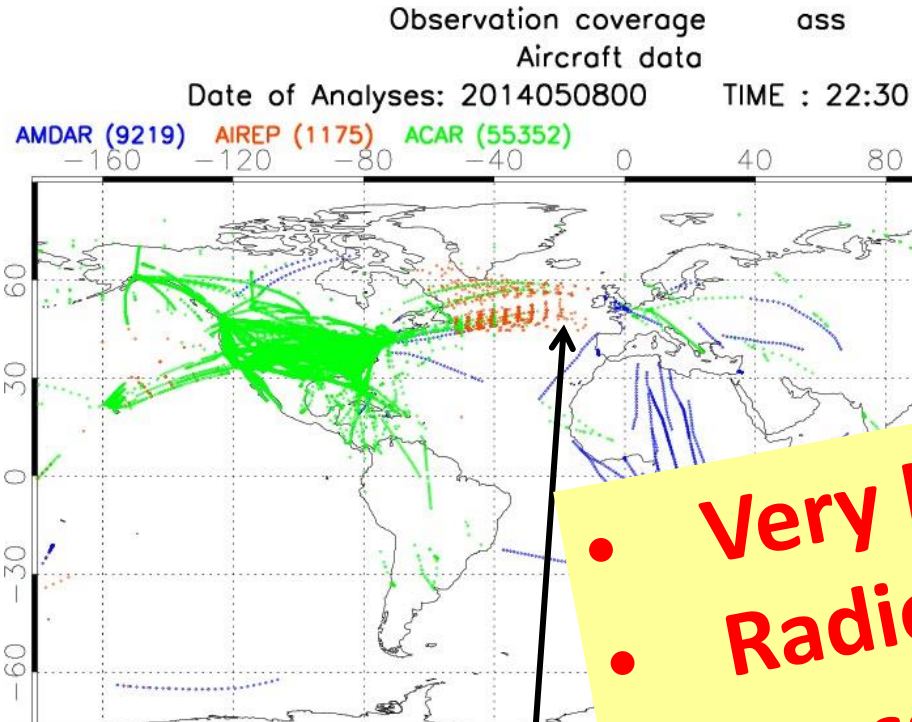
Radiosonde

Meteor. Observatorium Lindenberg

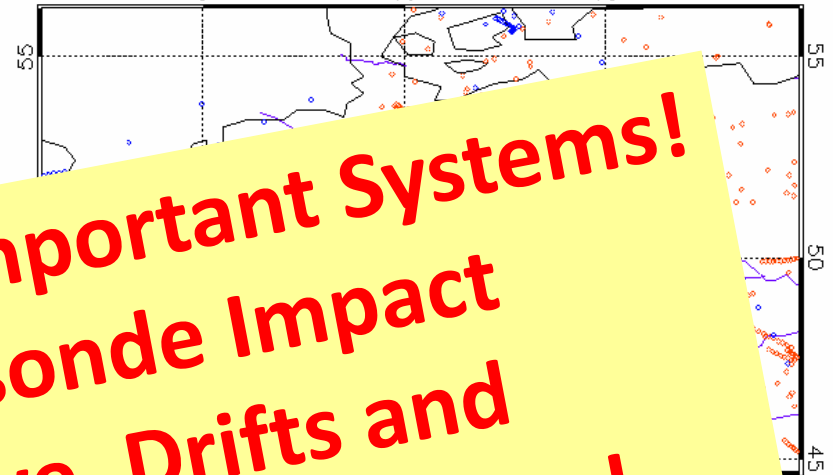
1. Doppler-LIDAR (Wind)
2. DIAL (Humidity)
3. Raman LIDAR (Temp+Hum)
4. MWR (Temp+Hum)
5. GPS STD (Hum)
6. Cloud Radar



Conventional Synop + Airplanes



Observation coverage - ass
Aircraft data
Date of Analyses: 20141101 TIME : 00:00 - 23:59
AMDAR (279) AIREP (0) ACAR (1284)

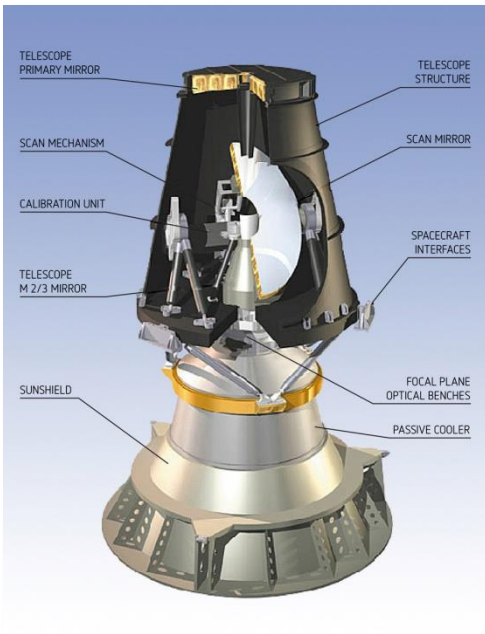


- Very Important Systems!
- Radiosonde Impact massive. Drifts and High-Resolution Used.
- Mode-S Operational over Germany (KENDA)

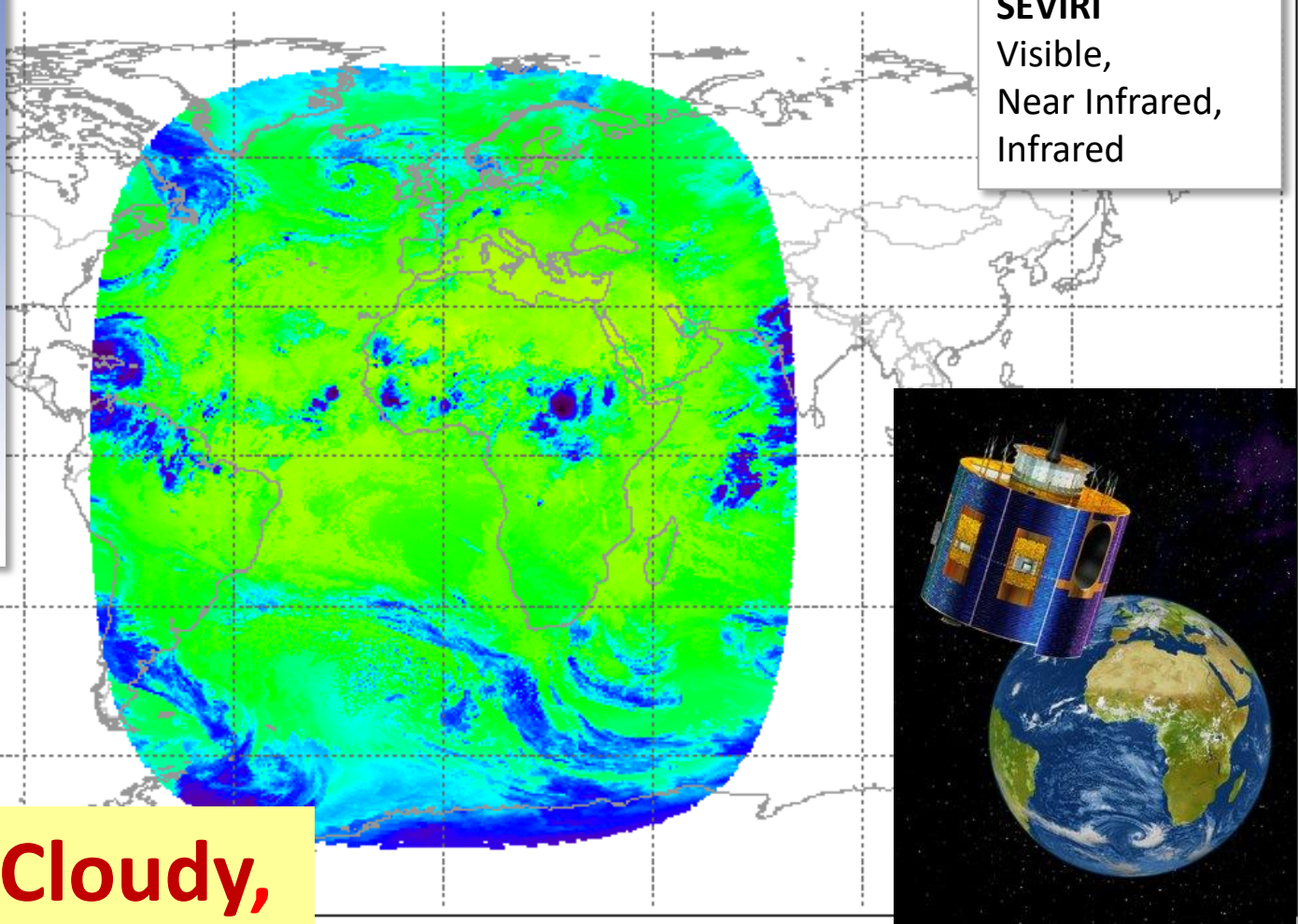


Observations: Geostationary Satellites

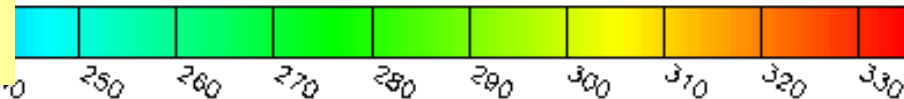
Bild: Robin Faulwetter



SEVIRI
Visible,
Near Infrared,
Infrared



**Clear and Cloudy,
IR, VIS (MFASIS)**



Observations: Polar Orbiting Satellites

Deutscher Wetterdienst
Wetter und Klima aus einer Hand



MW and IR Sounder
and Imager
Research Work:
Interchannel Corr +
Over Land + All-Sky

Aeolus Wind Lidar: Cal-Val +
Assimilation tests

Hyperspectral:
PC Scores Assimilation

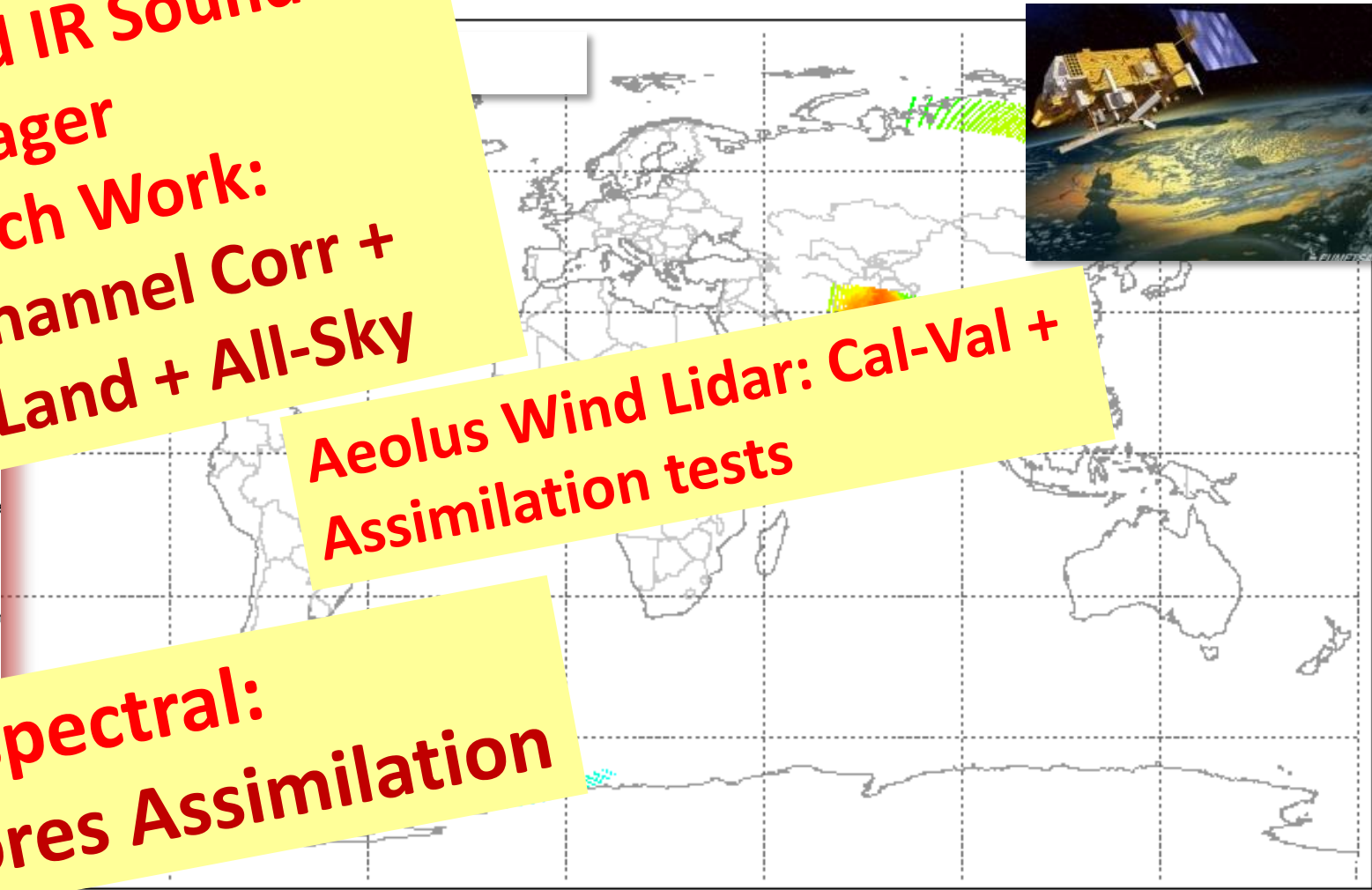
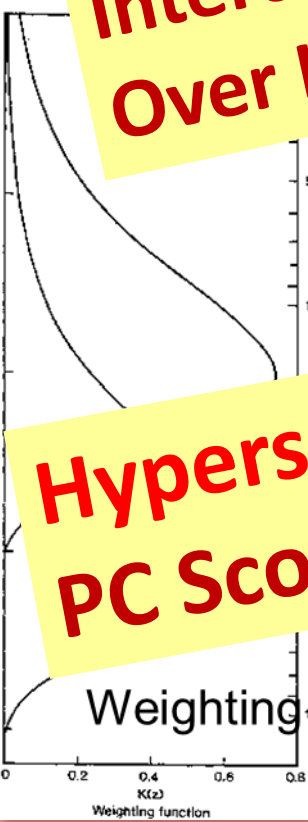
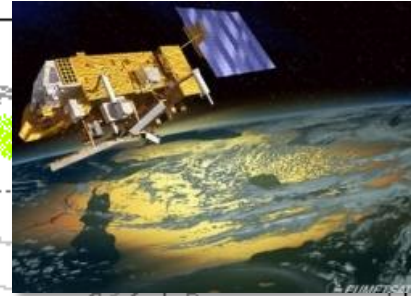


Bild: Robin Faulwetter

Ensemble Datenassimilation EnVar

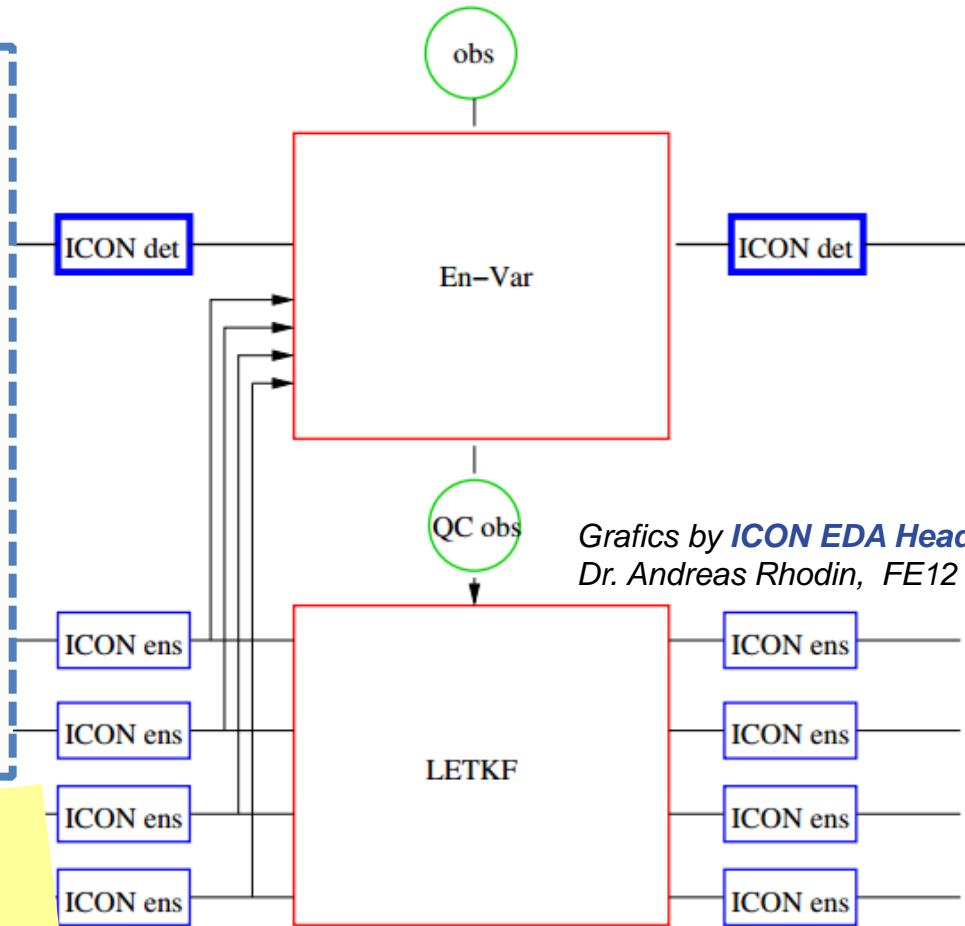
Operational since January 2016

Deutscher Wetterdienst
Wetter und Klima aus einer Hand



We are running **ICON EDA** in our Routine since Jan 2016

- 40 Members each with 40km global resolution and 20km NEST over Europe
- 1 deterministic 13km/6.5km
- **EPS forecasts** 40 Members 7 Days + 1 Deterministic
- Output for convective-scale EDA/EPS
- **Hybrid System**



Hybrid Ensemble - Variational, 3h cycle

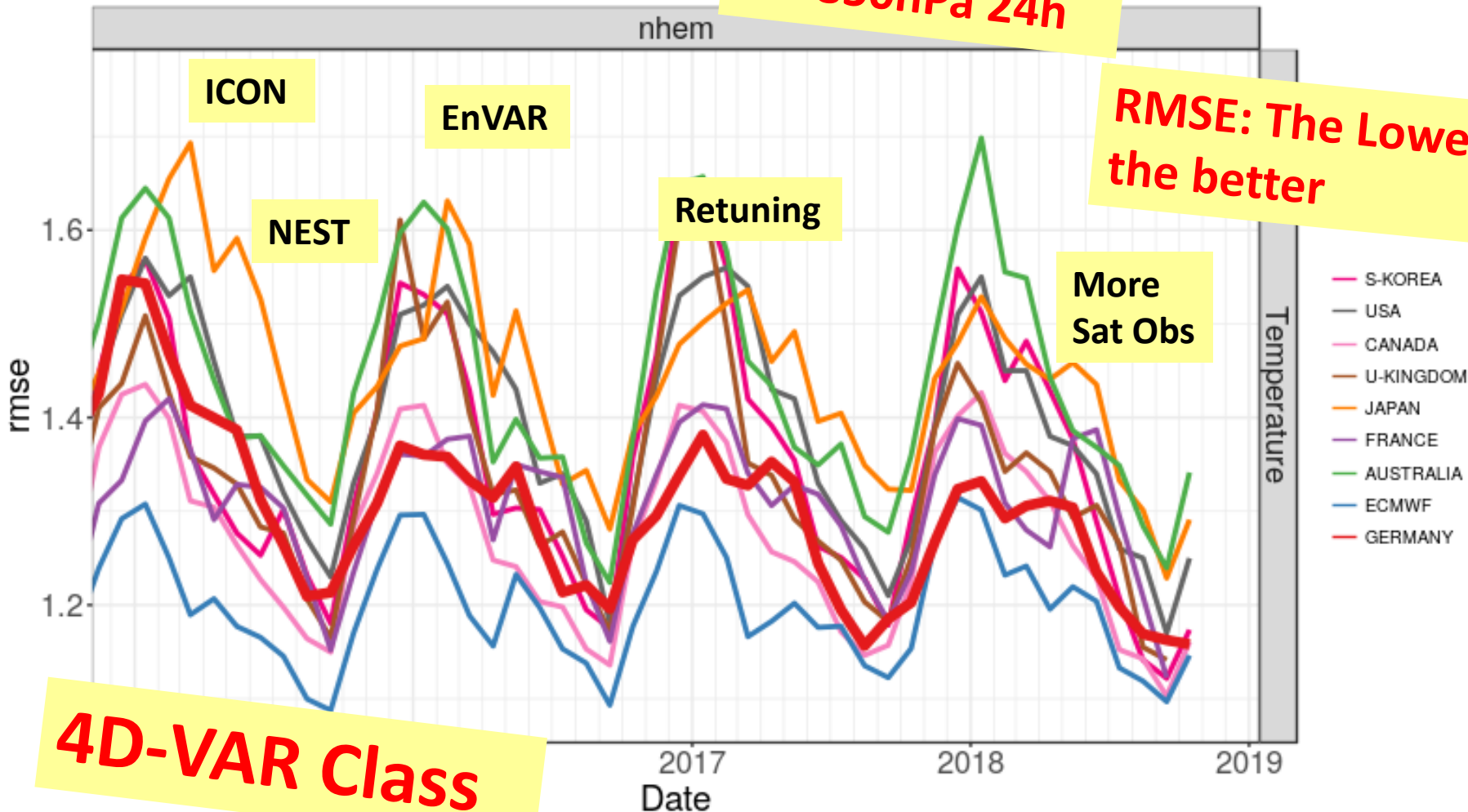


Hybrid Methods: EnVAR Scores

WMO verification against observations
lead-time: 24h
valid-time: 12UTC
level: 850hPa

Temperature
forecast quality
NH 850hPa 24h

RMSE: The Lower
the better



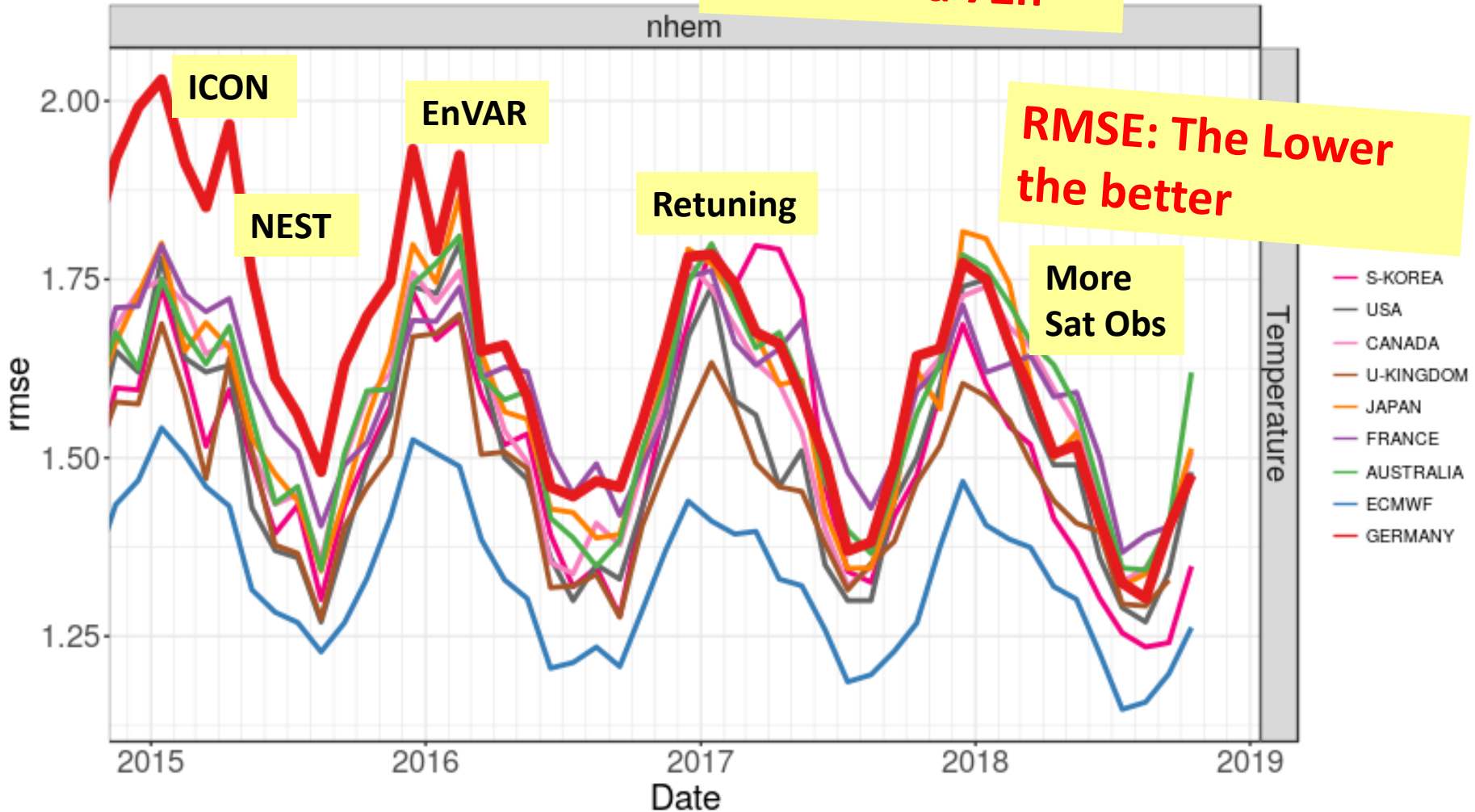
4D-VAR Class

Hybrid Methods: EnVAR Scores



Temperature
forecast quality,
NH 500hPa 72h

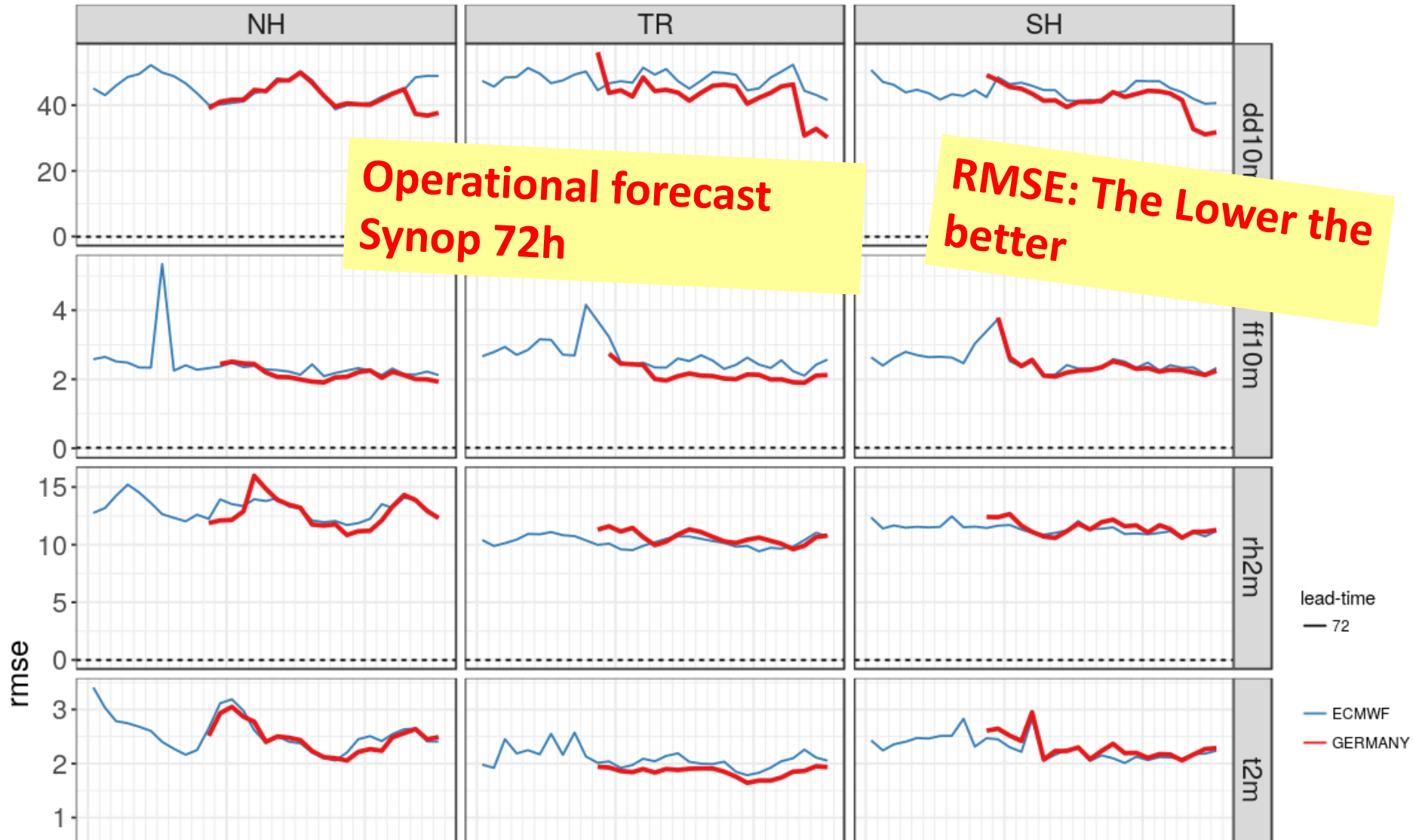
WMO verification against observations
lead-time: 72h
valid-time: 12UTC
level: 500hPa



Hybrid Methods: EnVAR Scores

WMO verification against SYNOP
lead-time: 72h
valid-time: 12UTC

2017+2018

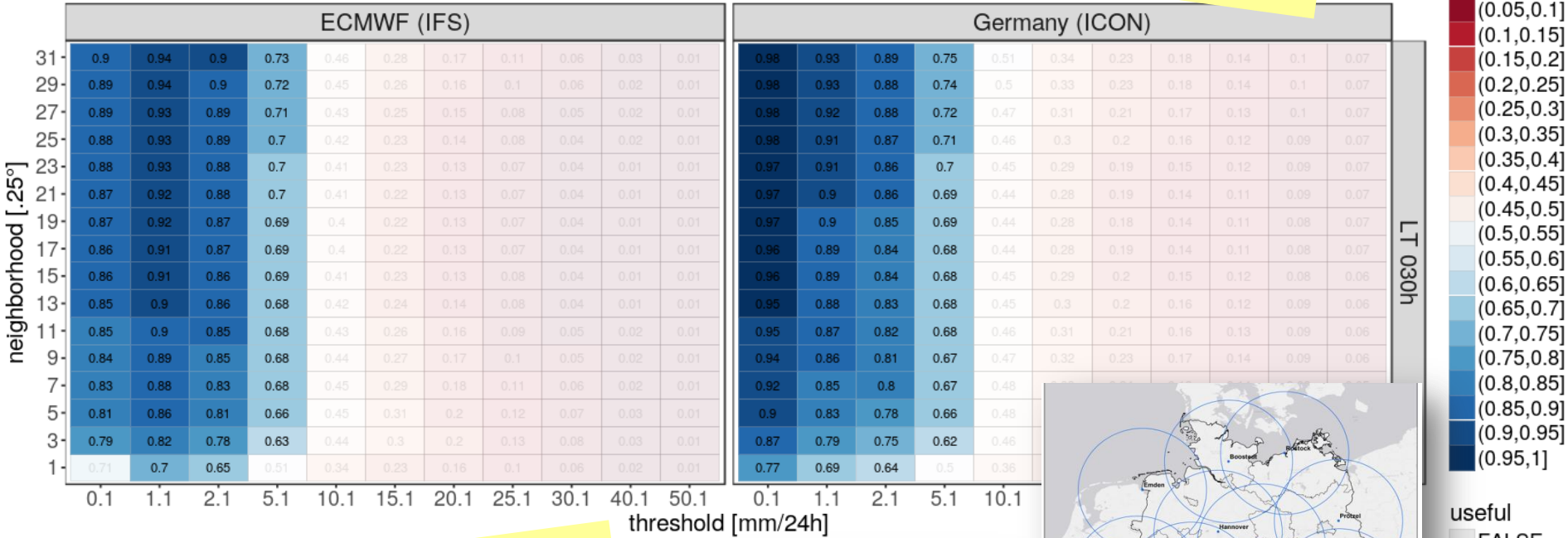


Hybrid Methods: EnVAR Scores

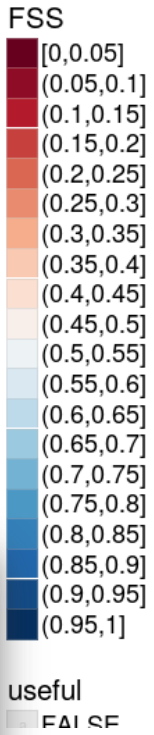
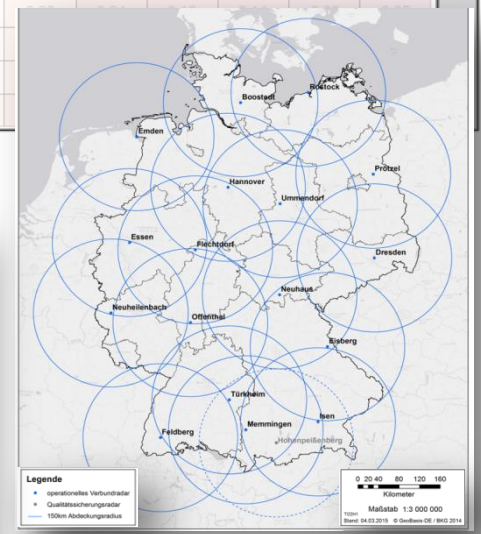
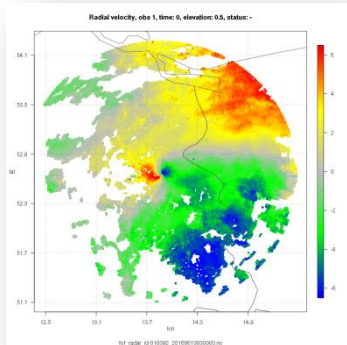
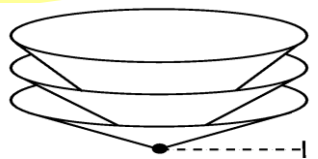


FSS: the higher the better

OBS: german radar (rw) 06UTC-06UTC
 PRED: rr_24h (0.25°)
 INI: 00UTC
 CASES: ECMWF (IFS) 90, Germany (ICON) 90



**Operational forecast
Precipitation vs RADAR**



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PRIOR

DATA

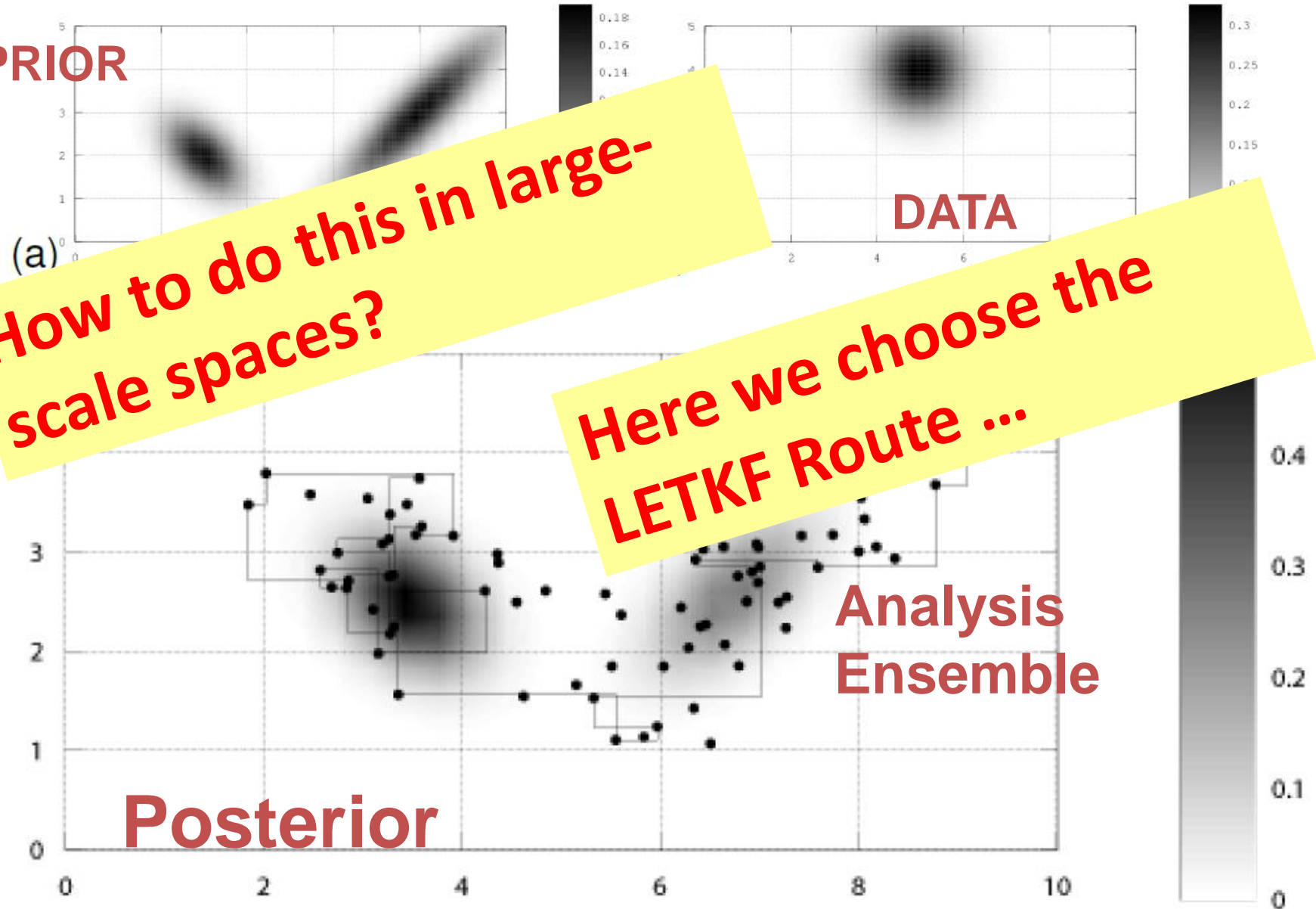
How to do this in large-scale spaces?

Here we choose the LETKF Route ...

Analysis Ensemble

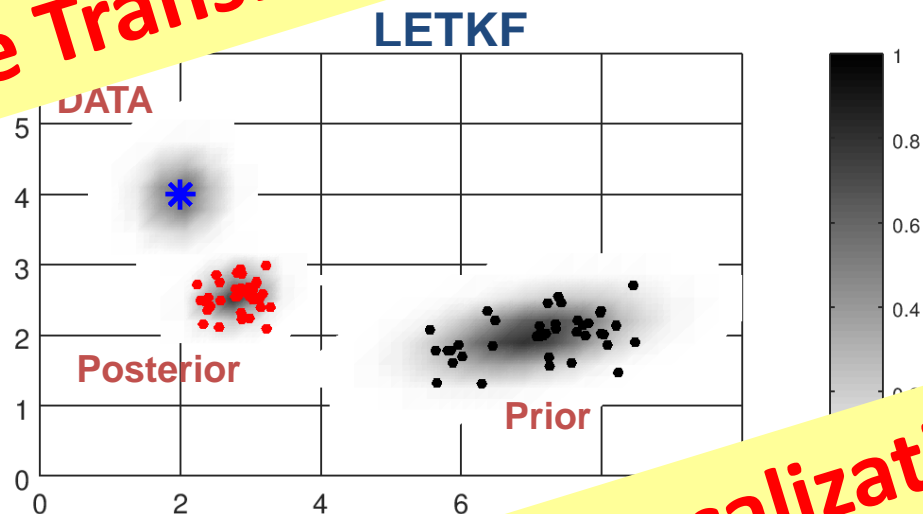
Posterior

(c)



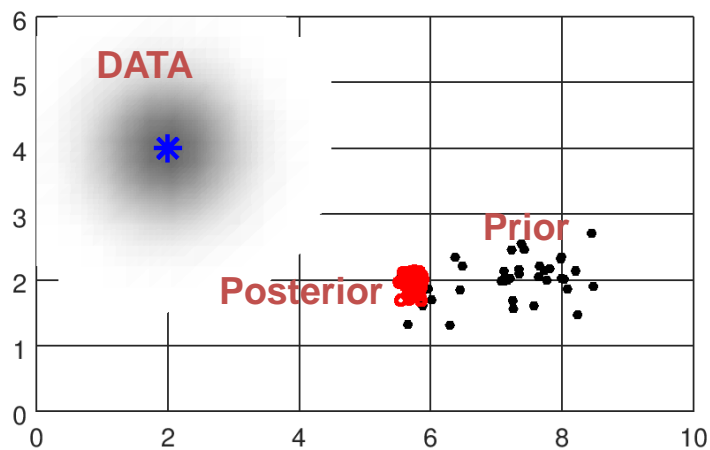
Bayesian Filtering via PF

1) Ensemble Transform

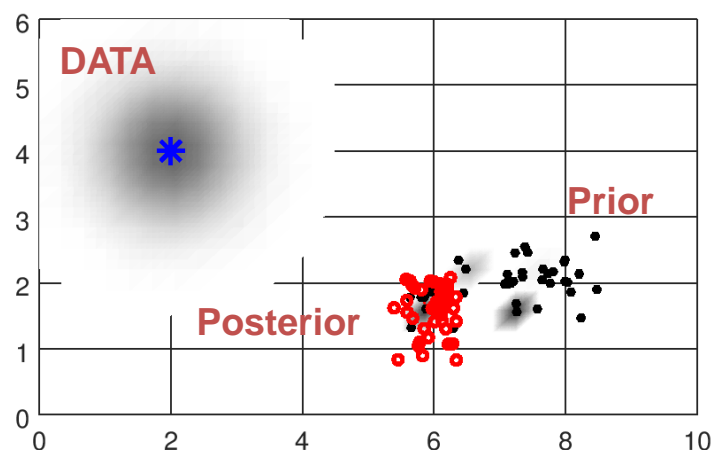


2) Localization

Classical PF



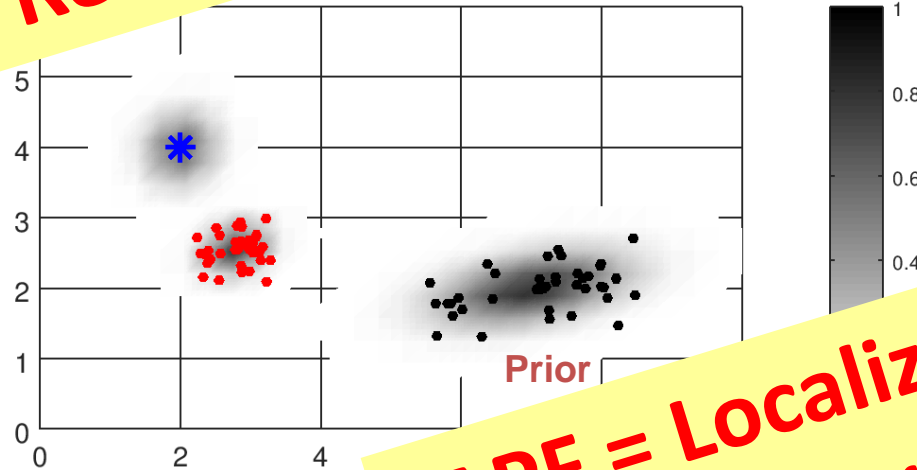
LAPF



Bayesian Filtering via PF

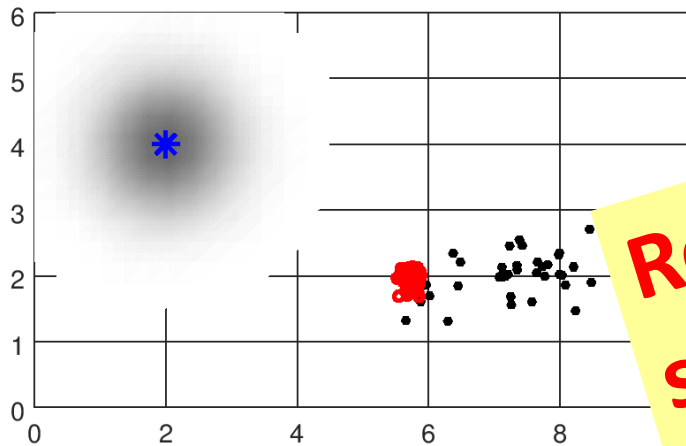
3) Adaptive Resampling

LETKF



LAPF = Localized
Adaptive Particle Filter

Classical PF



LAPF

Resampling is in ensemble
space, global + modulated

LAPF = Transform, Localization, Adaptivity with global modulated Resampling

- **Bayes formula** to calculate new analysis distribution

$$p_k^{(a)}(x) := p(x|y_k) = c p(y_k|x) p_k^{(b)}(x), \quad x \in \mathbb{R}^n$$

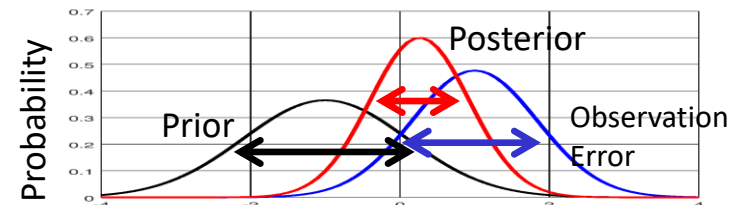
c is a normalization factor: $\int_X p_k^{(a)}(x) dx = 1$

Classical PF Approach

- To carry out the analysis step at time t_k
aposteriori weights $p_k^{(a)}$ are calculated

$$p_{k,l}^{(a)} = c e^{-\frac{1}{2}(y-Hx^{(l)})^T R^{-1}(y-Hx^{(l)})}$$

c is chosen such that $\sum_{l=1}^L p_{k,l}^{(a)} = L$



- **Accumulated weights** w_{ac} are defined:

$$w_{ac_0} = 0$$
$$w_{ac_i} = w_{ac_{i-1}} + p_i^a, \quad i = 1, \dots, L$$

where L denotes the ensemble size

- Drawing $r_j \sim U([0,1])$, $j = 1, \dots, L$, set $R_j = j - 1 + r_j$ and define **transform matrix** W for the particles by:

$$W_{i,j} = \begin{cases} 1 & \text{if } R_j \in (w_{ac_{i-1}}, w_{ac_i}], \\ 0 & \text{otherwise,} \end{cases}$$

$i, j = 1, \dots, L$ with $W \in \mathbb{R}^{L \times L}$, $(s, t]$ denotes the interval of values $s < \eta \leq t$.

Resampling

Adaptivity based on o-b statistics

- Based on the adaptive multiplicative **inflation factor** ρ determined by the LETKF

$$\rho = \frac{\text{E} \left[\mathbf{d}_{o-b}^T \mathbf{d}_{o-b} \right] - \text{Tr}(\mathbf{R})}{\text{Tr}(\mathbf{H}\mathbf{P}^b\mathbf{H}^T)}$$

from $\text{E} \left[\mathbf{d}_{o-b}^T \mathbf{d}_{o-b} \right] = \text{Tr}(\mathbf{R}) + \rho \text{Tr}(\mathbf{H}\mathbf{P}^b\mathbf{H}^T)$

- Weighting factor** α has been chosen, due to the small ensemble size ($L = 40$)

$$\rho_k = \alpha \tilde{\rho}_k + (1 - \alpha) \rho_{k-1}$$

- **Perturbation factor σ** is used to add spread to the system

$$\sigma = \left\{ \begin{array}{ll} c_0, & \rho < \rho^{(0)} \\ c_0 + (c_1 - c_0) * \frac{\rho - \rho^{(0)}}{\rho^{(1)} - \rho^{(0)}}, & \rho^{(0)} \leq \rho \leq \rho^{(1)} \\ c_1, & \rho > \rho^{(1)} \end{array} \right\}$$

where $c_0 = 0.02$, $c_1 = 0.2$,

$\rho^{(0)} = 1.0$ and $\rho^{(1)} = 1.4$, with

$\sigma = c_1$ if $\rho \geq \rho^{(1)}$ and

$\sigma = c_0$ if $\rho \leq \rho^{(0)}$

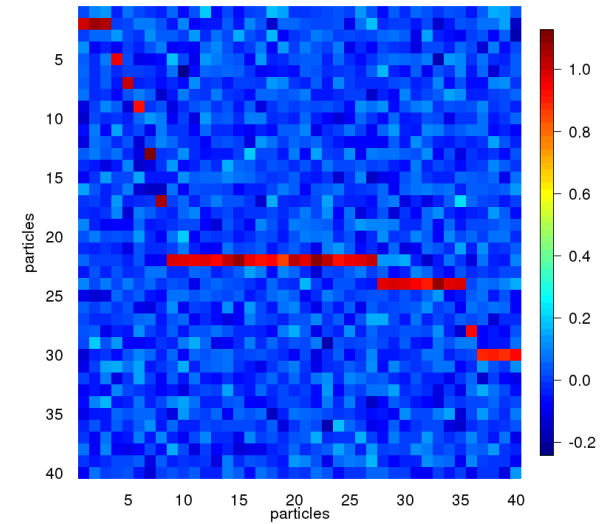
**Enforce the
desired spread!**

Fourth Step: Gaussian Resampling

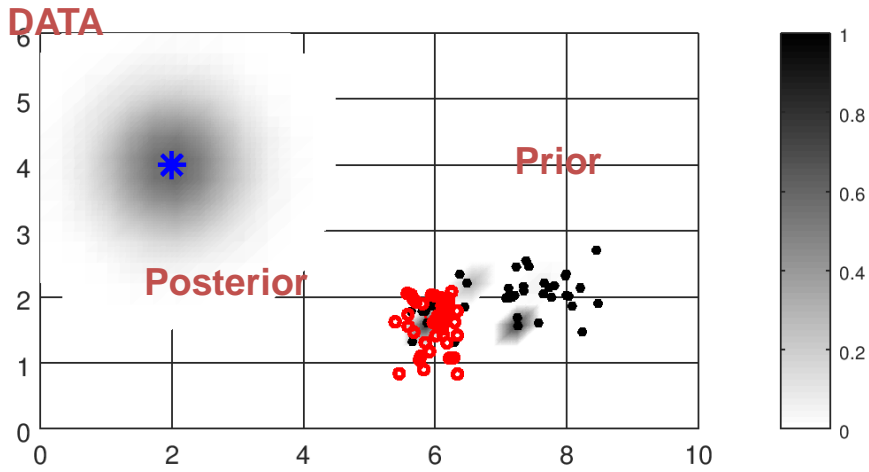
- Weights W are modified by applying the **perturbation factor σ**

$$W = W + R_{nd} * \sigma$$

with R_{nd} normally distributed random numbers



An example for a W-Matrix after applying σ determined with $\sigma = 0.5$ for 60%

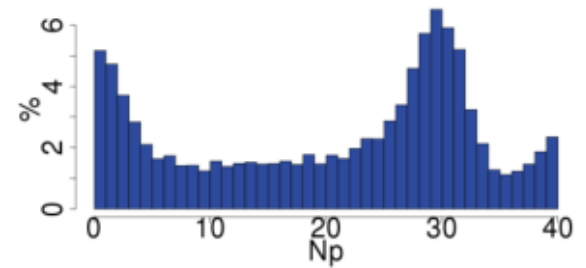
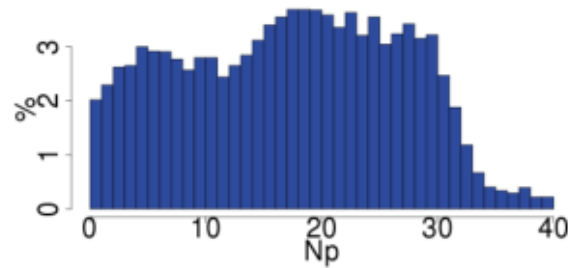
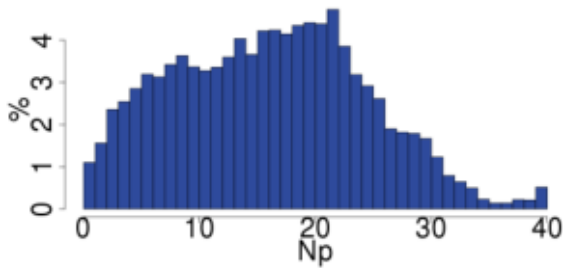


Enforce the desired spread!

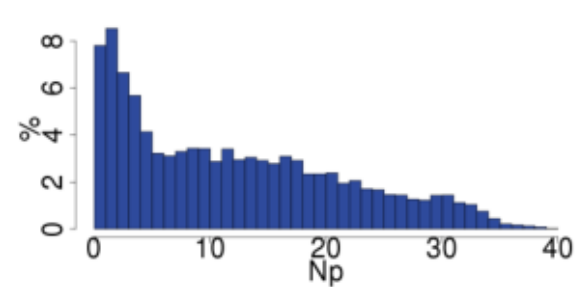
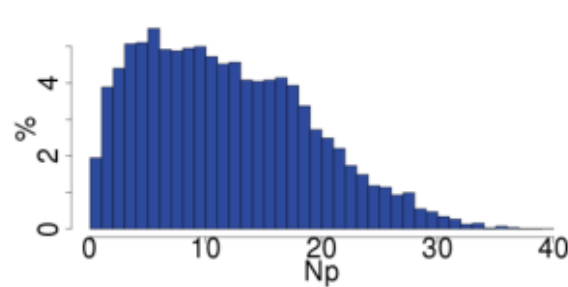
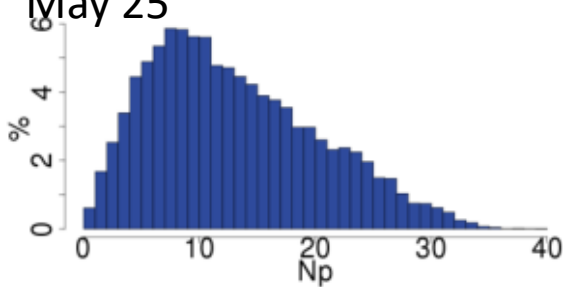
are chosen

Effective Ensemble Size Distributions

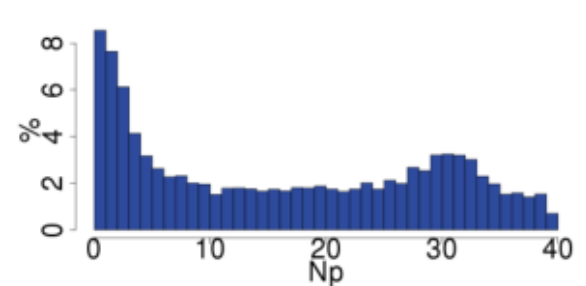
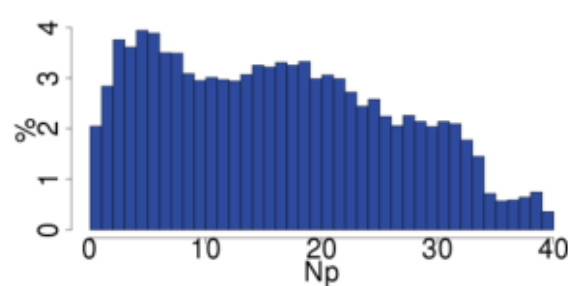
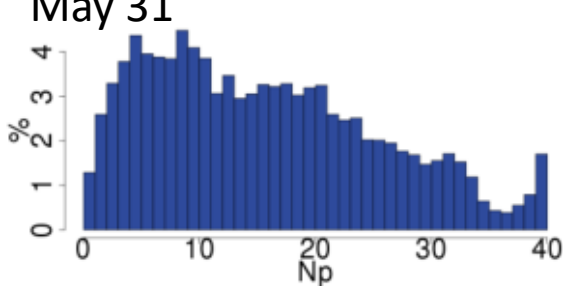
May 20



May 25



May 31



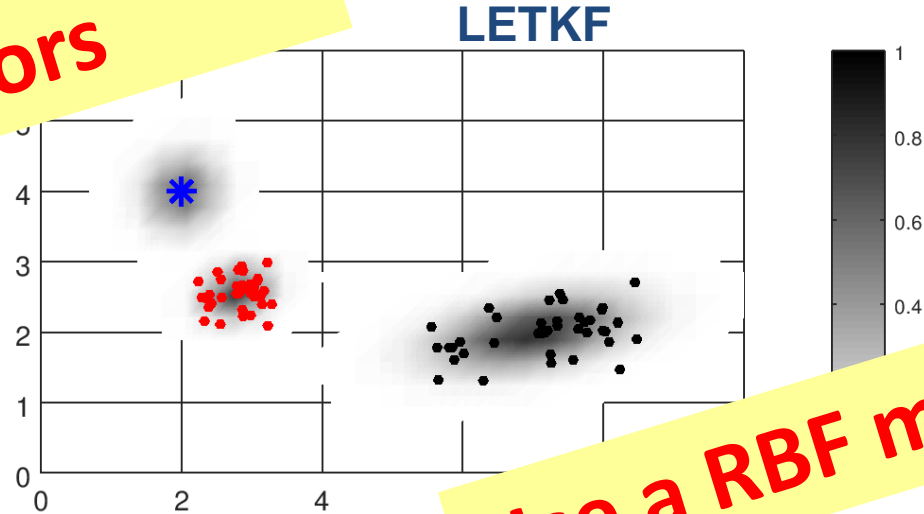
100 hPa

500 hPa

1000 hPa

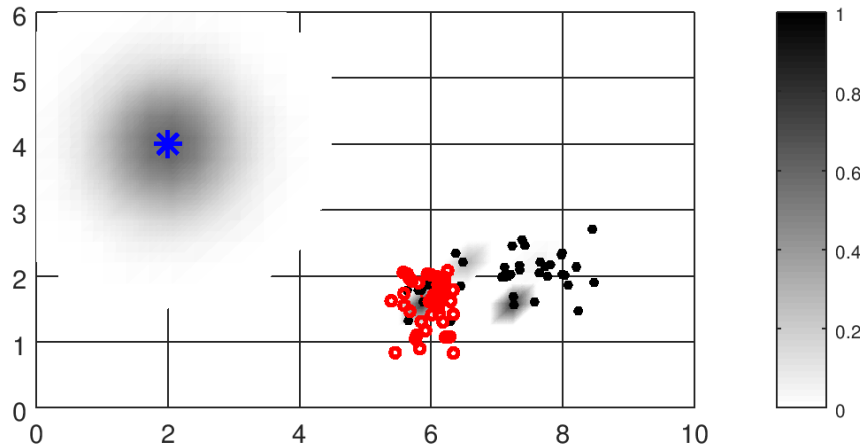
LAPF versus LMCPF

LAPF has problems
with model errors

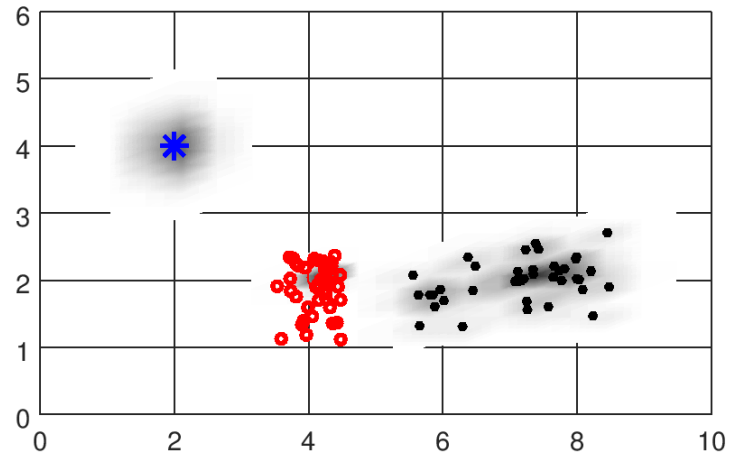


Use a RBF mixture

LAPF



LMCPF



LMCPF = Transform, Localization, RBF mixture, Adaptivity

■ Kalman Filter

$$x^{(a)} = x^{(b)} + BH^T(R + HBH^T)^{-1}(y - Hx^{(b)})$$

$$K = BH^T(R + HBH^T)^{-1} \quad \tilde{B} = (I - KH)B$$

■ Ensemble B Estimator

$$\bar{x} := \frac{1}{L} \sum_{\ell=1}^L x^{(\ell)}$$

$$B = \frac{1}{L-1} X X^T$$

$$X = (x^{(1)} - \bar{x}, \dots, x^{(L)} - \bar{x}) \in \mathbb{R}^{n \times L}.$$

$$\begin{aligned}\tilde{B} &= (I - KH)B \\ &= (I - BH^T(R + HBH^T)^{-1}H)B \\ &= \left(I - \gamma X \boxed{X^T H^T} (R + \gamma \boxed{HX} \boxed{X^T H^T})^{-1} H \right) \gamma X X^T \\ &= X \left(I - \gamma Y^T (R + \gamma Y Y^T)^{-1} Y \right) \gamma X^T \\ &= X \left(I - \gamma \boxed{(I + \gamma Y^T R^{-1} Y)^{-1} Y^T R^{-1} Y} \right) \gamma X^T \\ &= X \left((I + \gamma Y^T R^{-1} Y)^{-1} \boxed{(I + \gamma Y^T R^{-1} Y)} - \gamma Y^T R^{-1} Y \right) \gamma X^T \\ &= X (I + \gamma Y^T R^{-1} Y)^{-1} \gamma X^T \\ &= X \left(\frac{1}{\gamma} I + Y^T R^{-1} Y \right)^{-1} X^T\end{aligned}$$

$$Y := HX$$

**RBF Basis Function
in Ensemble Space**

$$Y^T (R + \gamma Y Y^T)^{-1} = (I + \gamma Y^T R^{-1} Y)^{-1} Y^T R^{-1}$$

$$p(x|y) = cp(y|x) \cdot p(x)$$

Gaussian Mixture Case

$$= ce^{-\frac{1}{2}(y-Hx)^T R^{-1}(y-Hx)} \sum_{\ell=1}^L p_G(x - x^{(\ell)})$$

$$= c \sum_{\ell=1}^L e^{-\frac{1}{2}(y-Hx)^T R^{-1}(y-Hx)} p_G(x - x^{(\ell)}),$$

$$p_G(x - x^{(\ell)}) = \tilde{c} e^{-\frac{1}{2}(x-x^{(\ell)})^T G^{-1}(x-x^{(\ell)})}$$

**Explicit Calculations
possible for each term**

**We need a
selection based on
relative weights!**

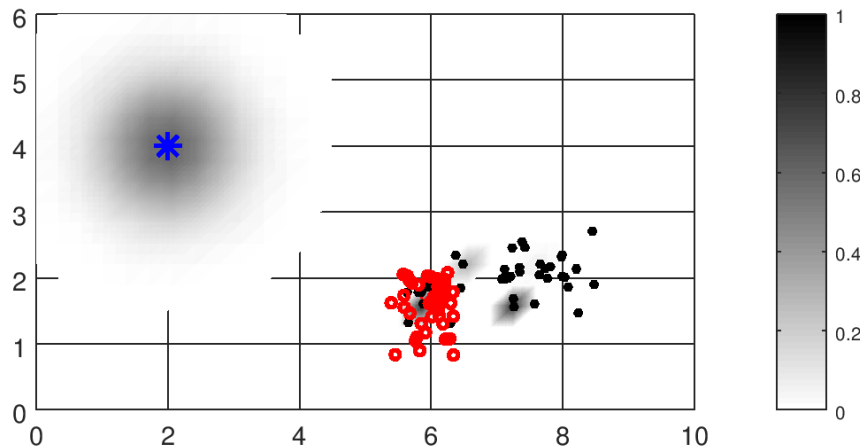
LMCPF Basics: Relative Weights

$$w_\ell := e^{-\frac{1}{2}(y-Hx^{(\ell)})^T R^{-1}(y-Hx^{(\ell)})}, \quad \ell = 1, \dots, L$$

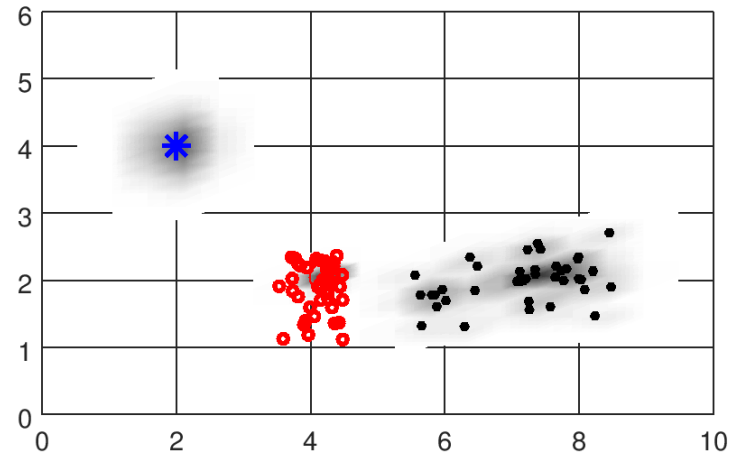
$$w_{tot} := \sum_{\ell=1}^L w_\ell.$$

We need a selection based on relative weights!

LAPF



LMCPF



Projection onto Ensemble Space

Abbreviating $A := \mathbf{Y}^T \mathbf{R}^{-1} \mathbf{Y}$ and $C := A^{-1} \mathbf{Y}^T \mathbf{R}^{-1} (\mathbf{y}^o - \bar{\mathbf{y}}^b)$

Projection Operator

$$P(\mathbf{y}^o - \bar{\mathbf{y}}^b) = \mathbf{Y}(\mathbf{Y}^T \mathbf{R}^{-1} \mathbf{Y})^{-1} \mathbf{Y}^T \mathbf{R}^{-1} (\mathbf{y}^o - \bar{\mathbf{y}}^b),$$

Projected discrepancy

$$P(\mathbf{y}^o - H\mathbf{x}^{(\ell)}) = \mathbf{Y}A^{-1}\mathbf{Y}^T\mathbf{R}^{-1}((\mathbf{y}^o - \bar{\mathbf{y}}^b) - \mathbf{Y}e_\ell)$$

Exponent

$$= \mathbf{Y}(C - e_\ell), \quad \ell = 1, \dots, L.$$

$$P(\mathbf{y}^o - H\mathbf{x}^{(\ell)})^T \mathbf{R}^{-1} P(\mathbf{y}^o - H\mathbf{x}^{(\ell)}) = [C - e_\ell]^T A [C - e_\ell], \quad \ell = 1, \dots, L,$$

Weight

$$w_{k,\ell} = ce^{-\frac{1}{2}[C - e_\ell]^T A [C - e_\ell]}, \quad \ell = 1, \dots, L.$$

Classical versus projected weights

$$\begin{aligned}w_{k,\ell}^{classical} &= e^{-\frac{1}{2}[(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]^T \mathbf{R}^{-1}[(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]} \\ &= e^{-\frac{1}{2}[(P+(I-P))(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]^T \mathbf{R}^{-1}[(P+(I-P))(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]} \\ &= e^{-\frac{1}{2}[P(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]^T \mathbf{R}^{-1}[P(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]} \cdot \underbrace{e^{-\frac{1}{2}[(I-P)(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]^T \mathbf{R}^{-1}[(I-P)(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]}}_{=\tilde{c}},\end{aligned}$$

Factor is a constant term, since we have

$$\begin{aligned}(I-P)(\mathbf{y}^o - H\mathbf{x}^{(\ell)}) &= (I-P)(\mathbf{y}^o - \bar{\mathbf{y}}^b + \mathbf{Y}e_\ell) \\ &= (I-P)(\mathbf{y}^o - \bar{\mathbf{y}}^b) - \underbrace{(I-P)\mathbf{Y}e_\ell}_{=0}.\end{aligned}$$

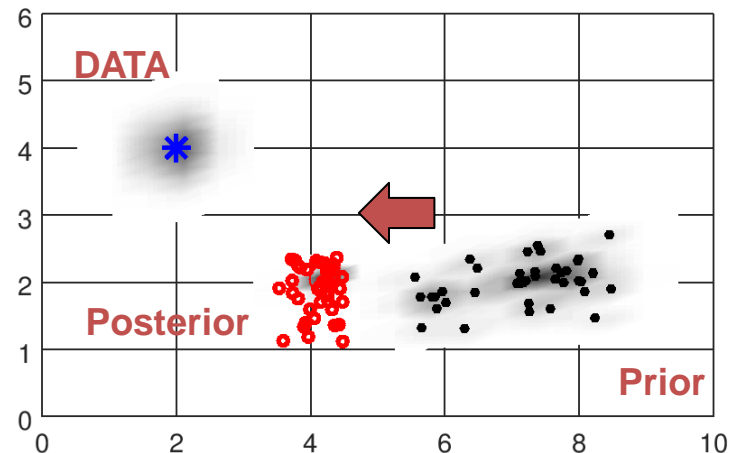
Projected particle filter weights and classical particle filter weights are equivalent theoretically, but numerically remove a very small common factor

LMCPF = Local Markov Chain Particle Filter

- **Weights W** are calculated by drawing from the posterior

$$W = W + A_{shift} * W + B_{post} * R_{nd} * \sigma$$

with R_{nd} normally distributed random numbers,
 A_{shift} and B_{post} calculated with Gaussian radial basis function (rbf) Approximation for prior density and observation error



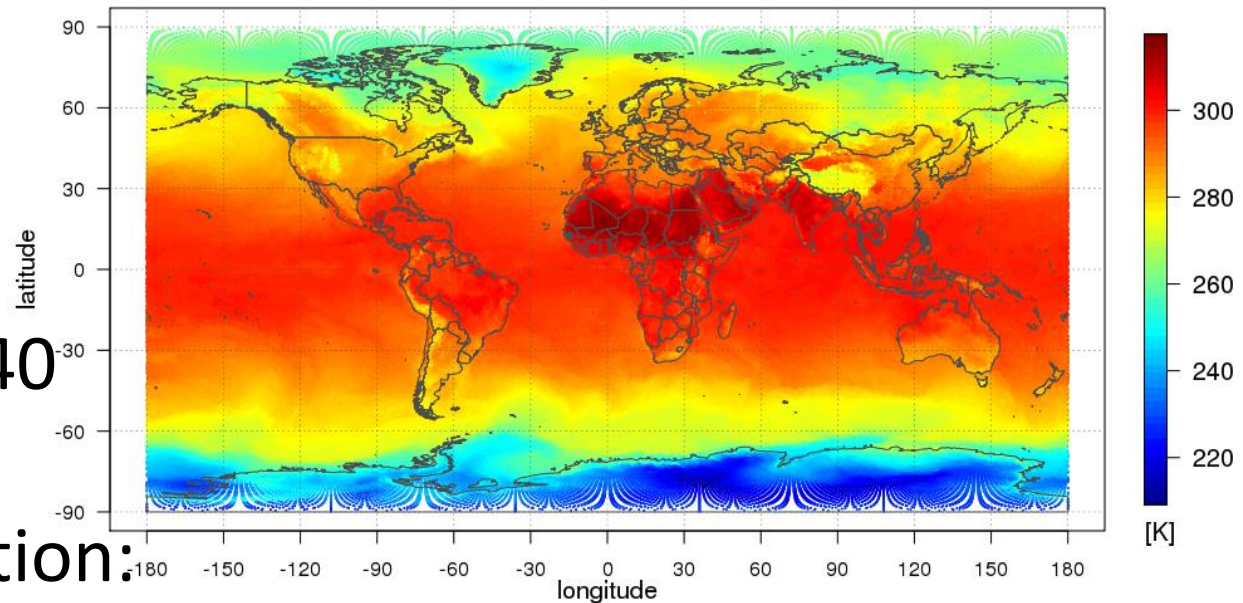
- ✓ It is an **explicit calculation of the Bayes posterior based on radial basis function approximation of the prior, with subsequent draws from that distribution in the MCMC sense.**

Large-Scale Experimental Set-up

Deutscher Wetterdienst
Wetter und Klima aus einer Hand



- Full ensemble: 40 members
- Reduced resolution:
 - 26km deterministic
 - 52km ensembles
- Period:
01.05.2016 –
31.05.2016

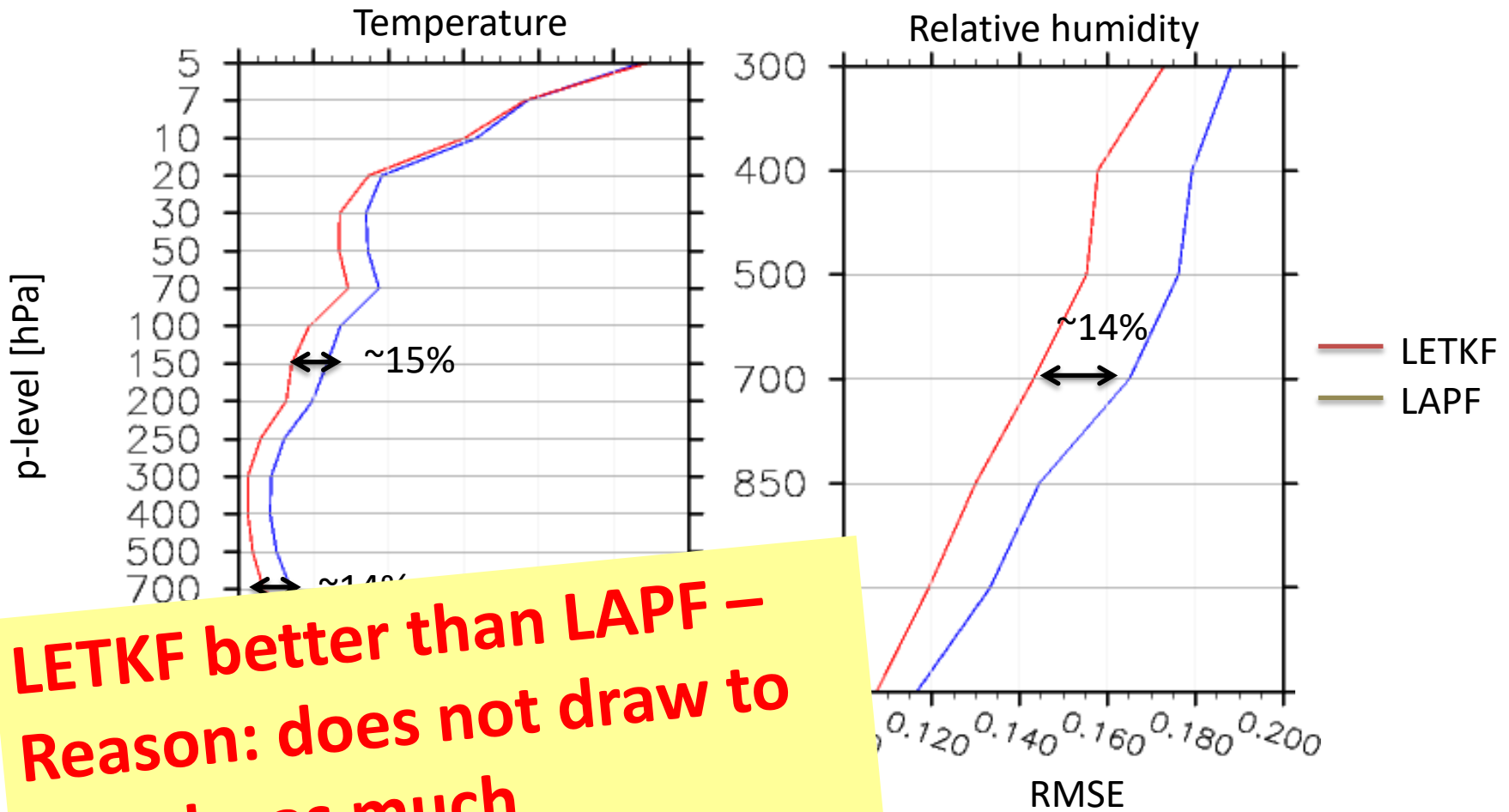


Experiments programmed and carried out
by **Anne Walter, DWD& Uni Reading**, and
Roland Potthast, DWD& Uni Reading

In Cooperation with Peter-Jan van
Leeuwen , Uni Reading

LAPF Scores vs LETKF

Global **RMSE** for **obs-fg** statistics (Radiosondes vs. Model)
Period: 08.05.2016 – 31.05.2016



**LETKF better than LAPF –
Reason: does not draw to
the obs as much**

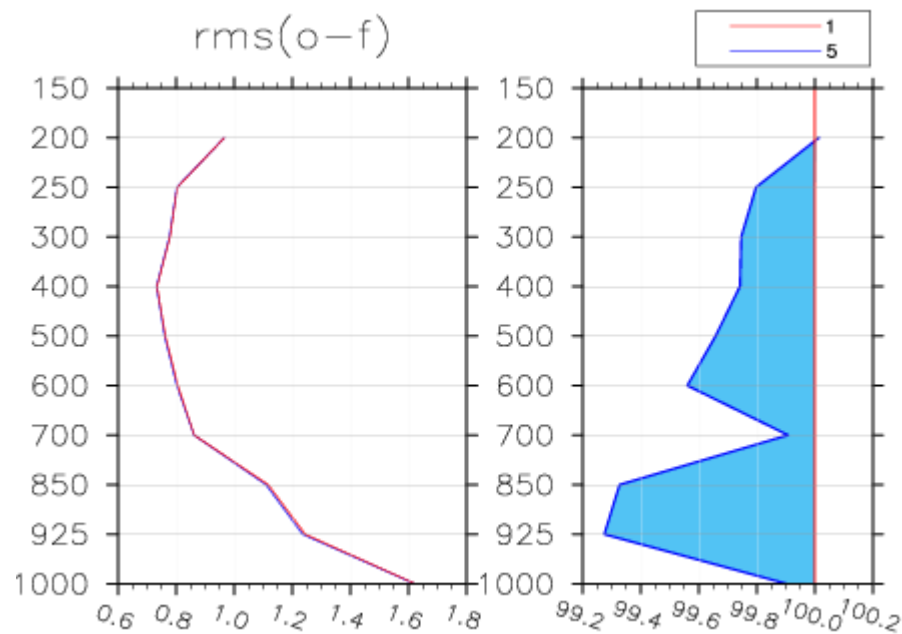
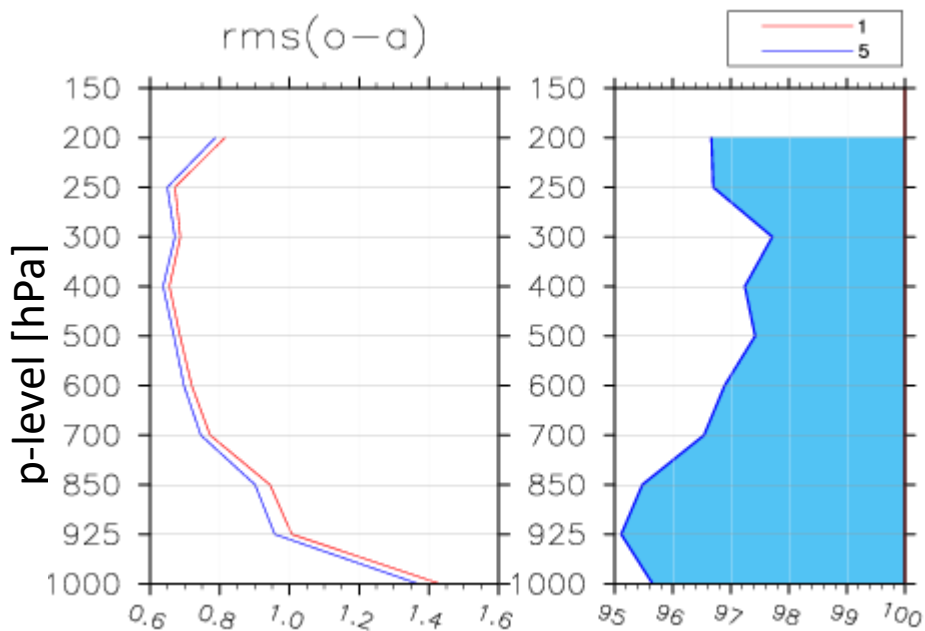
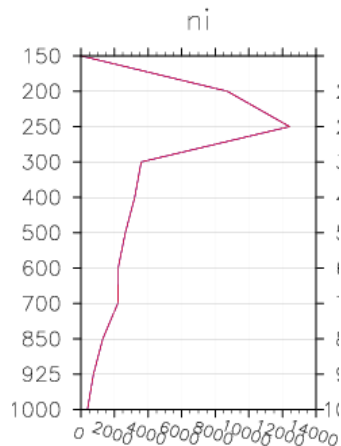
LMCPF Scores vs LETKF

Global **RMSE** for **obs-fg** statistics (AIREP vs. Model)
Period: 08.05.2016 – 22.05.2016

— LETKF
— LMCPF

Temperature o-a and o-f

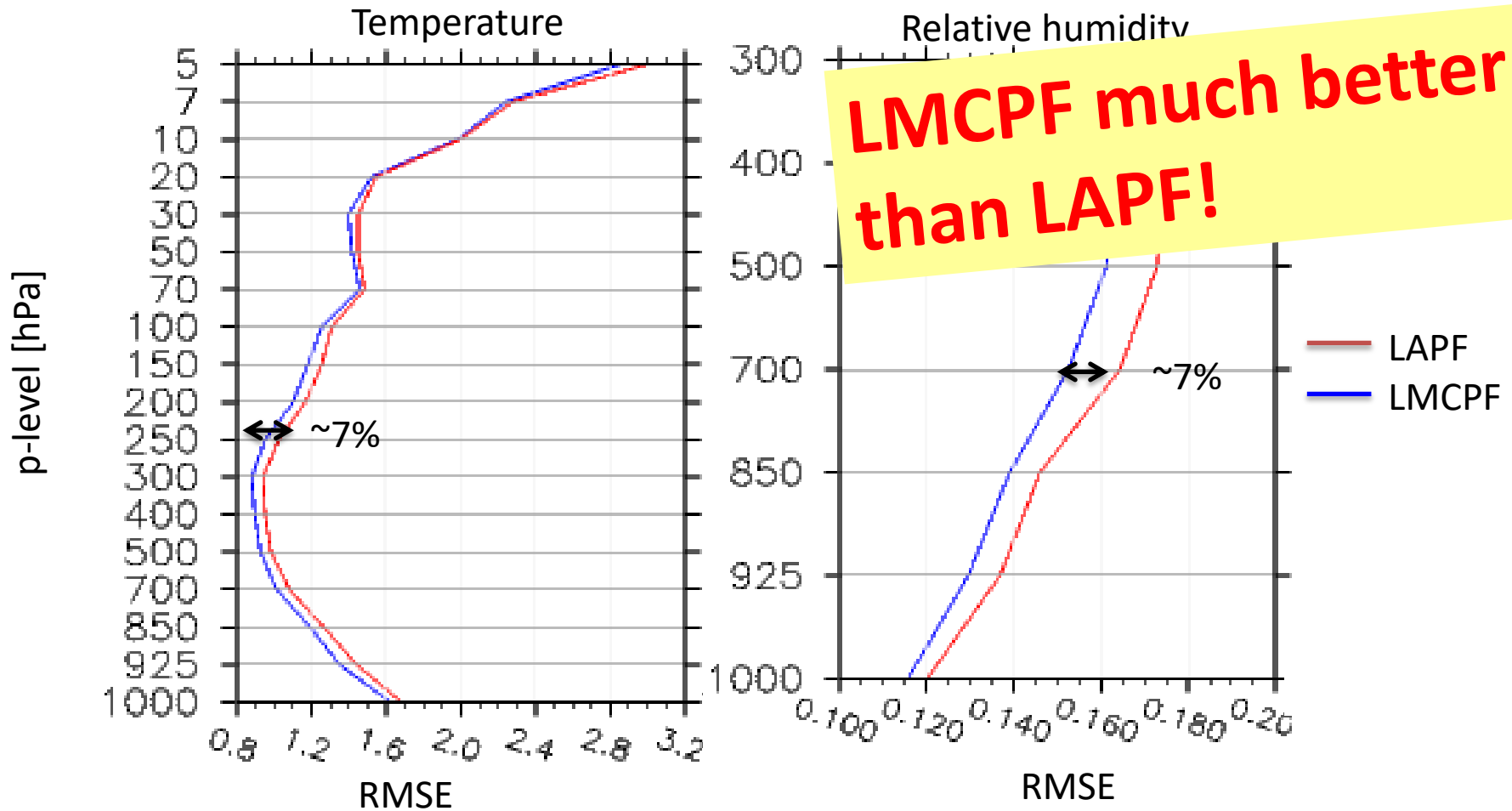
LMCPF: Model Error based Shift works



LMCPF Scores vs LAPF



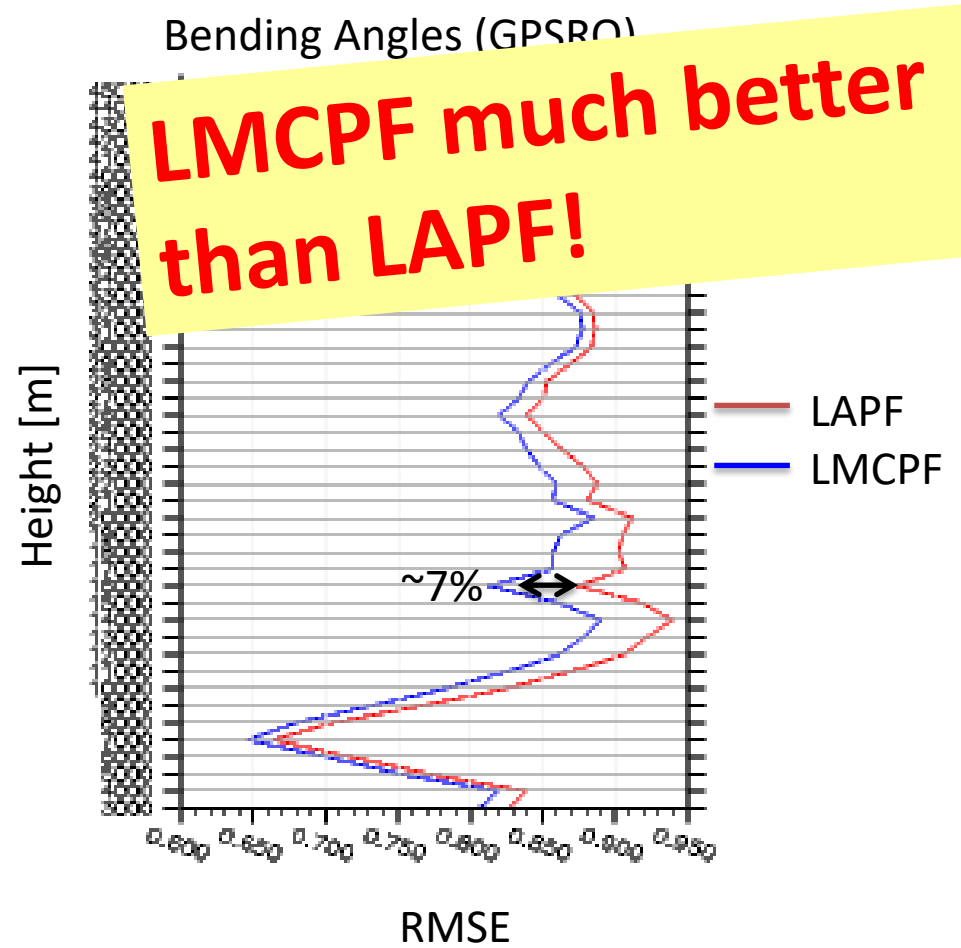
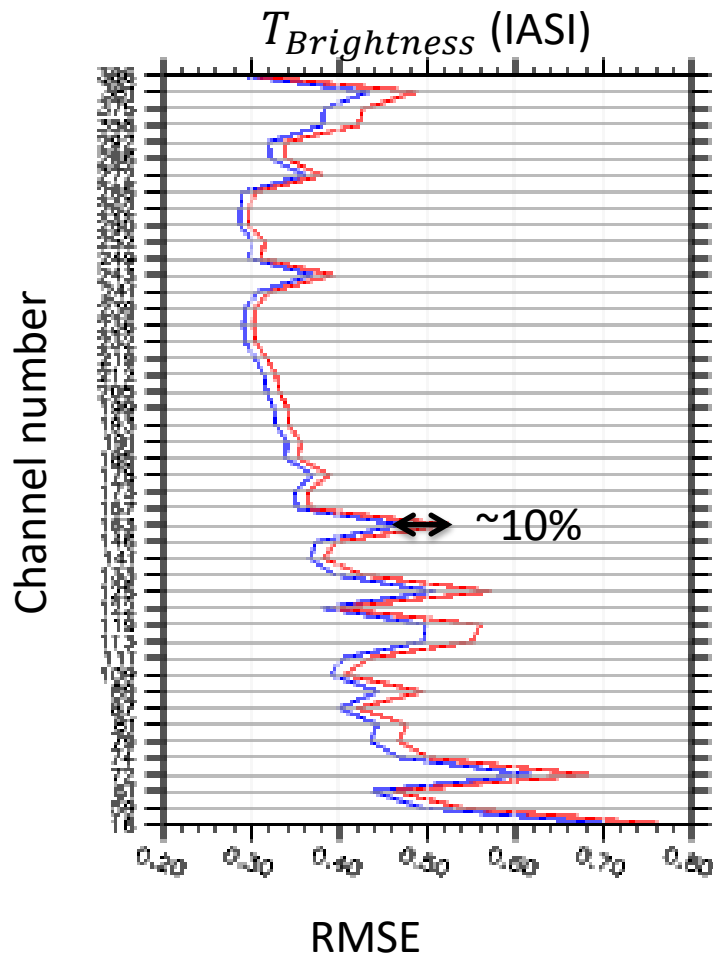
Global **RMSE** for **obs-fg** statistics (Radiosondes vs. Model)
Period: 08.05.2016 – 22.05.2016



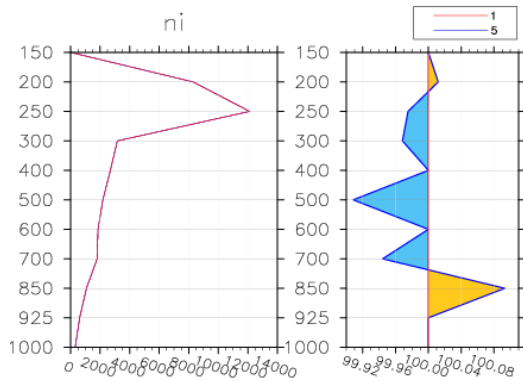
LMCPF Scores vs LAPF



Global **RMSE** for **obs-fg** statistics
Period: 08.05.2016 – 22.05.2016

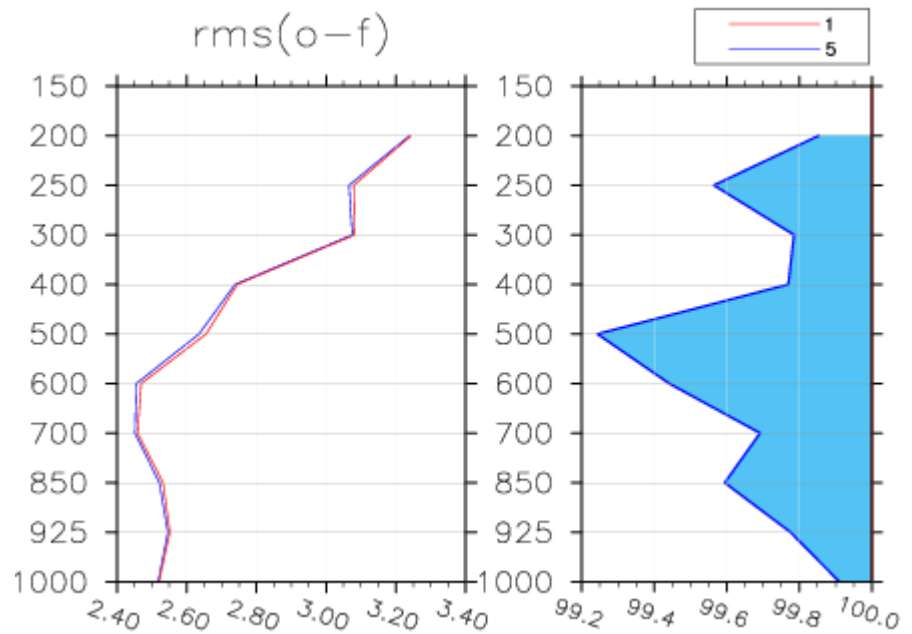
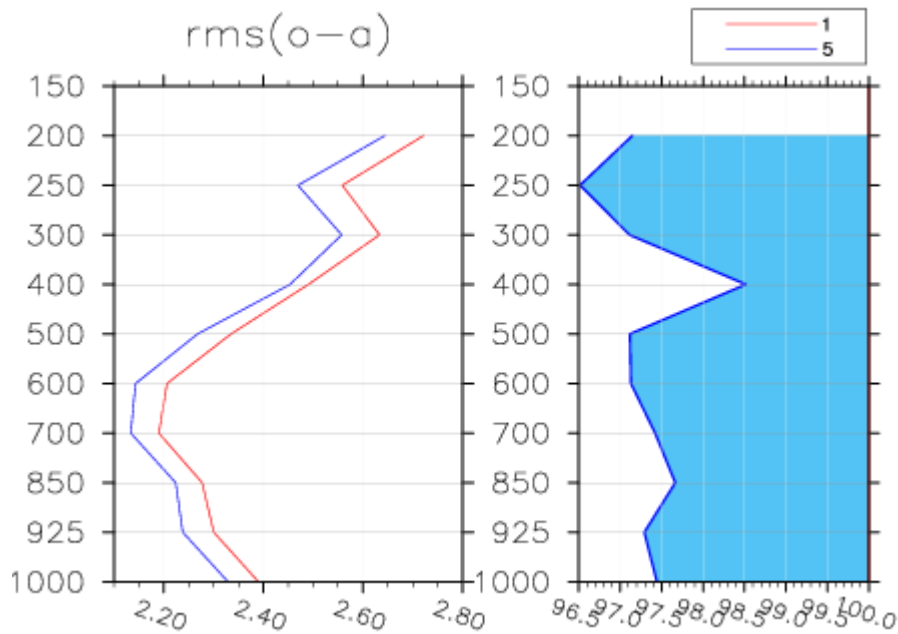


New LMCPF Scores vs LETKF



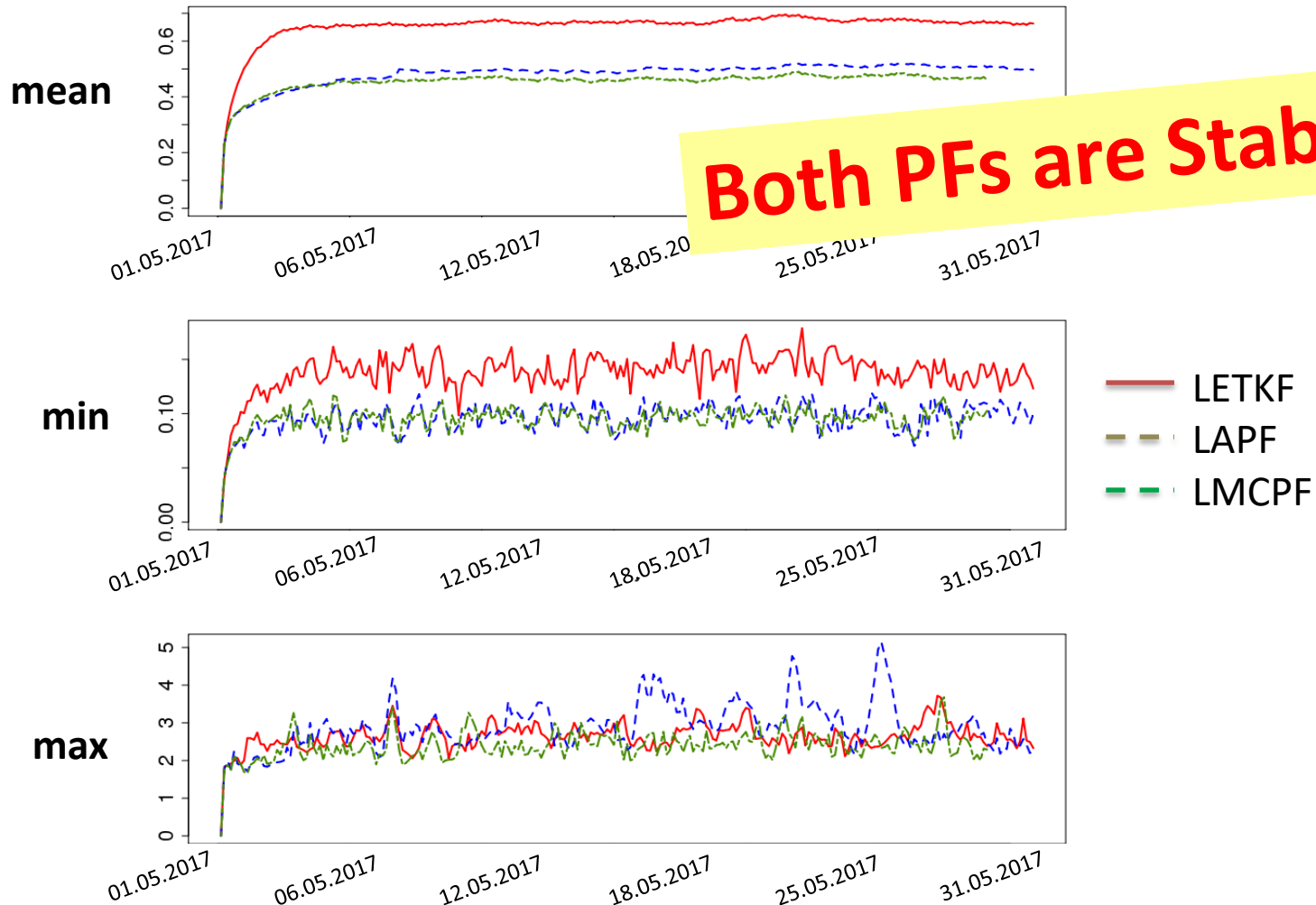
WIND o-a and o-f

LMCPF: Model Error based Shift works



LAPF Spread vs LMCPF & LETKF

Global spread of T [K] ~ 500 hPa



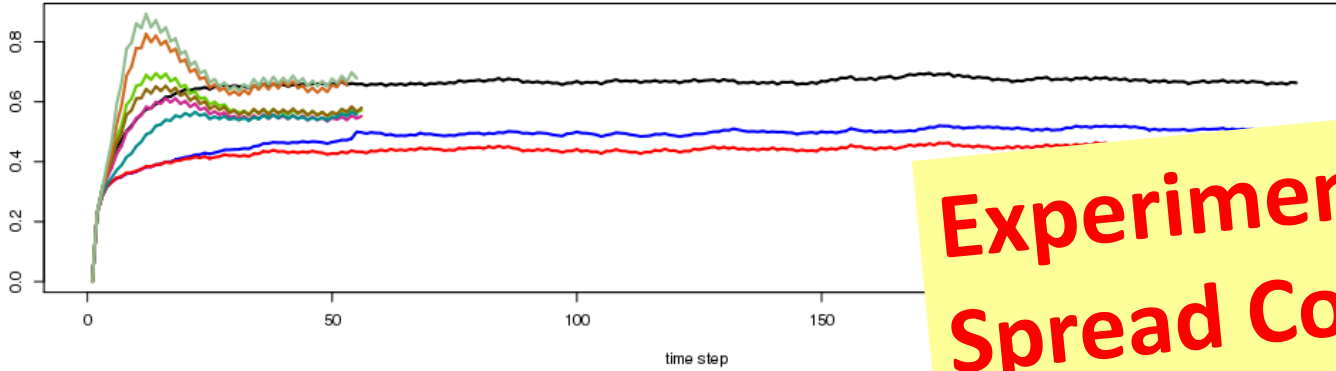
Both PFs are Stable!

Statistics for spread at level 64 for variable T



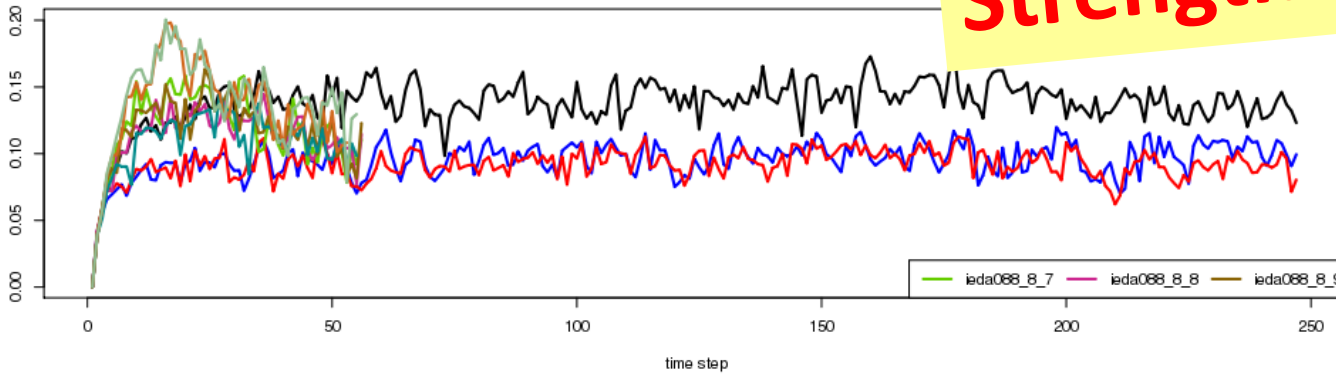
Deutscher Wetterdienst
Wetter und Klima aus einer Hand

Mean of spread

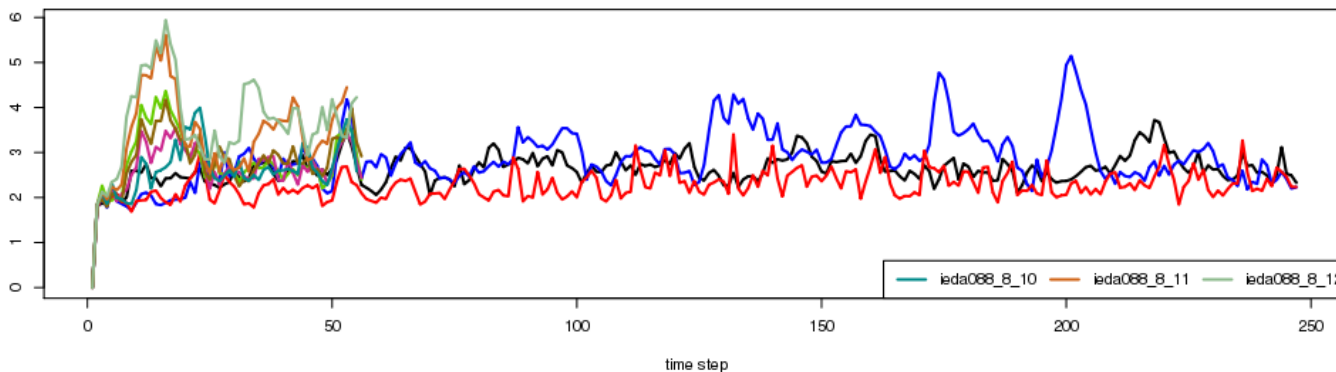


**Experiments with
Spread Control and
Strength of Shifts**

Minimum of spread



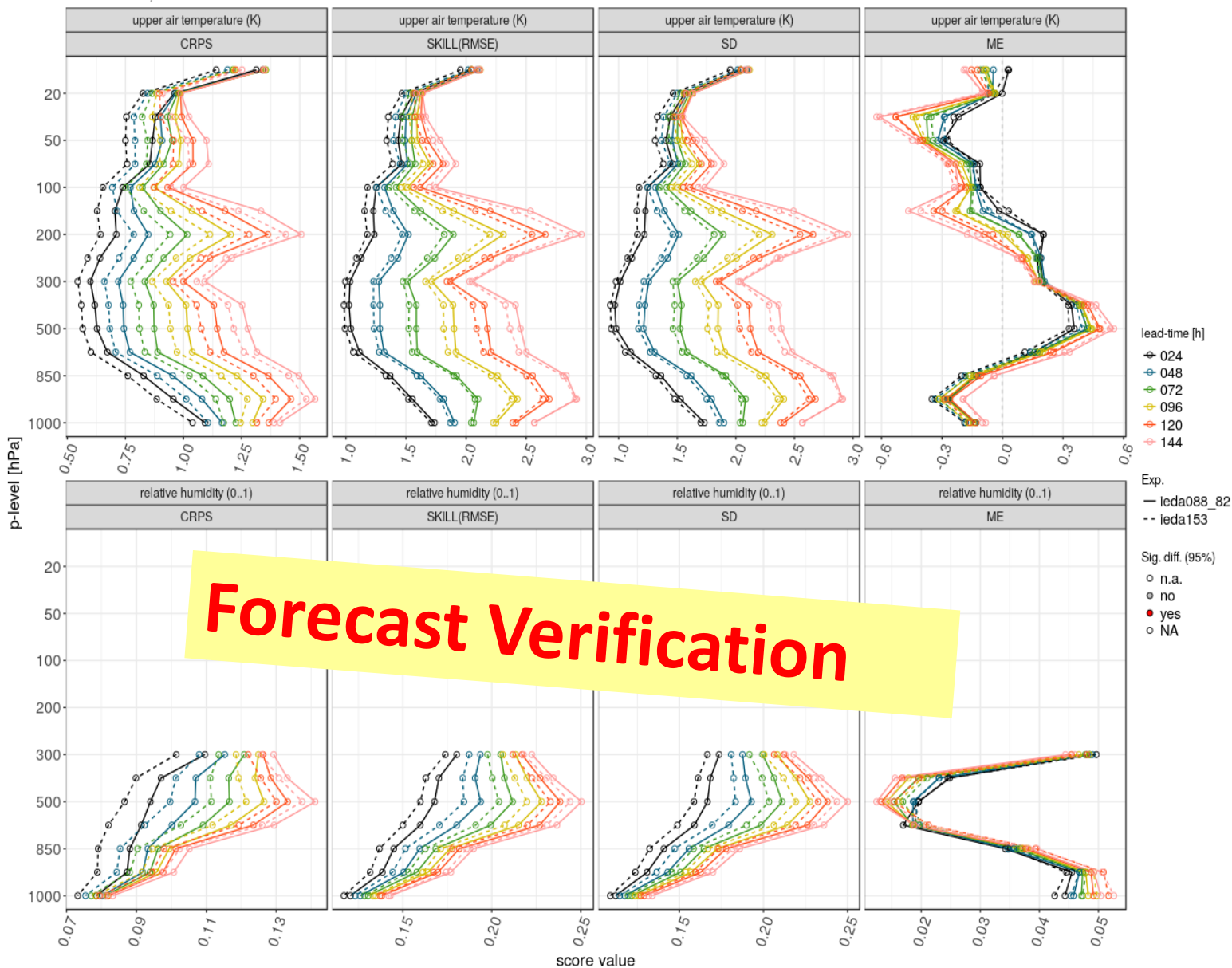
Maximum of spread



LMCPF Scores vs LETKF



2016/05/02 - 2016/05/24
INI: ALL UTC, DOM: ALL

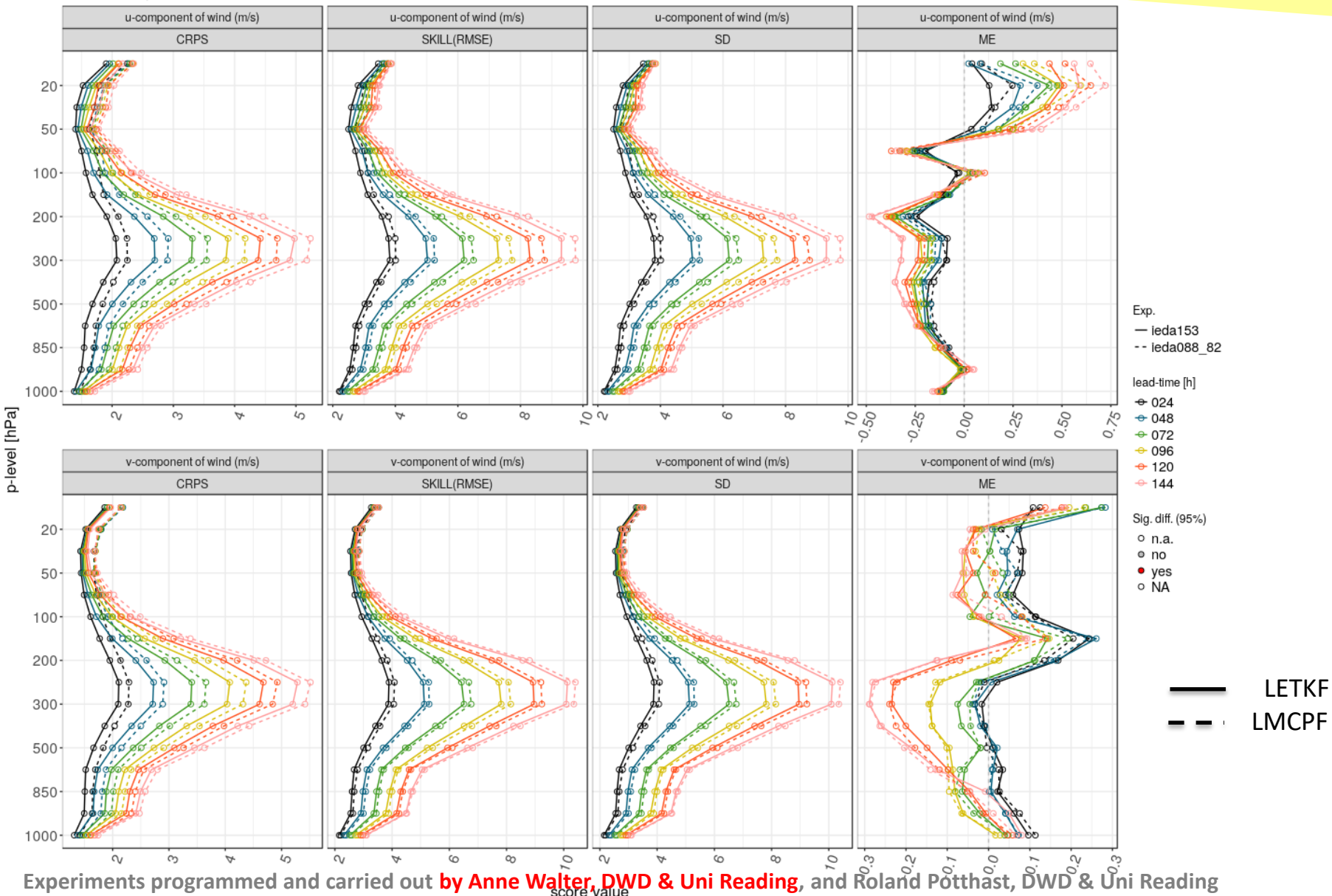


LMCPF Scores vs LETKF

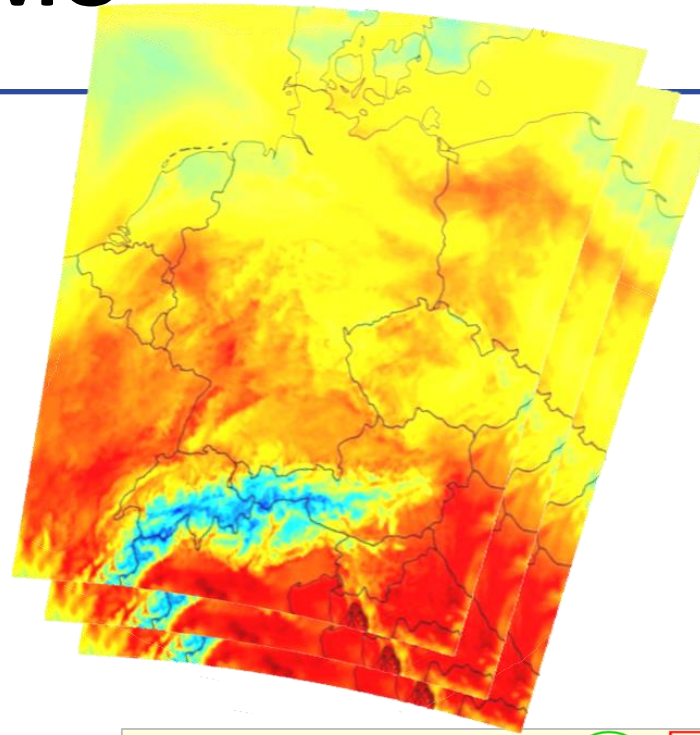


Forecast Verification

2016/05/02 - 2016/05/24
INI: ALL UTC, DOM: ALL

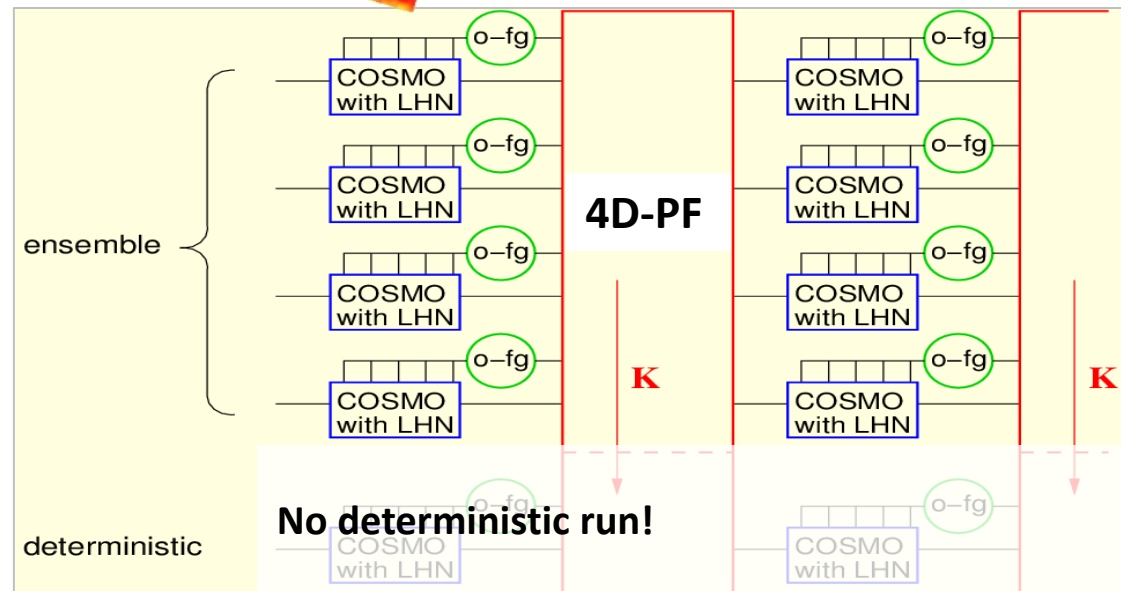


LMCPF for COSMO



4D-Particle Filter for Convection Permitting Model

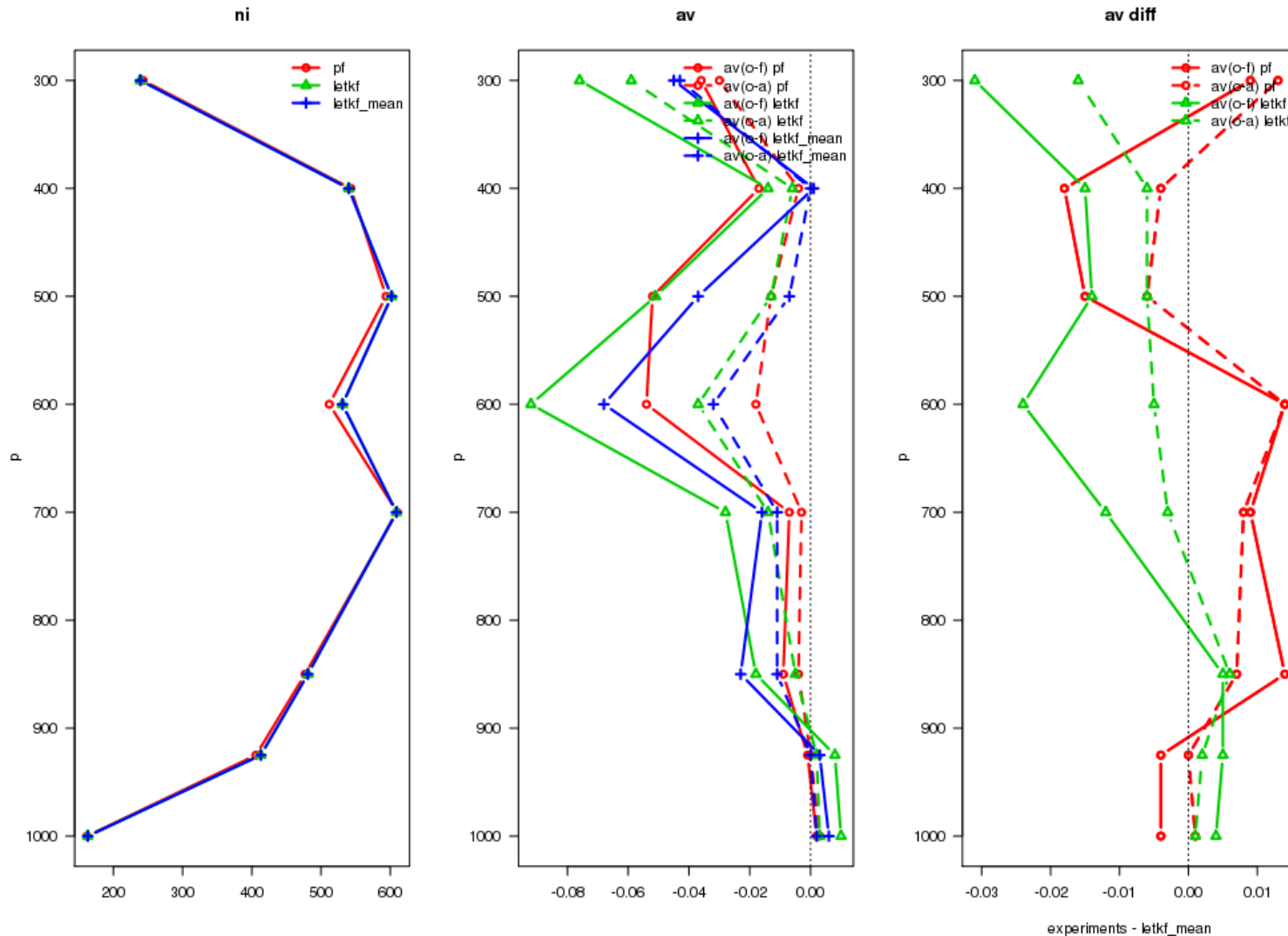
COSMO-DE
Resolution
2.8km
Central Europe



LMCPF for COSMO

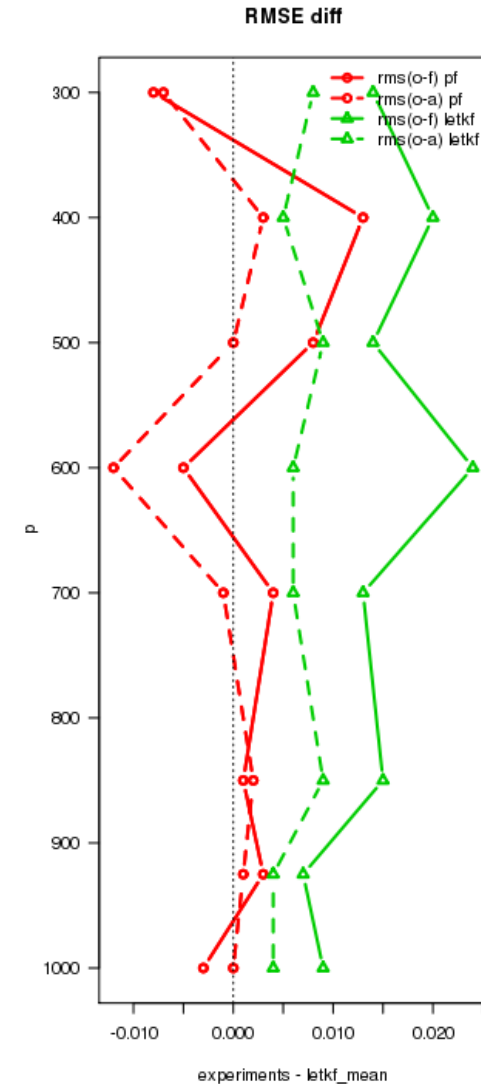
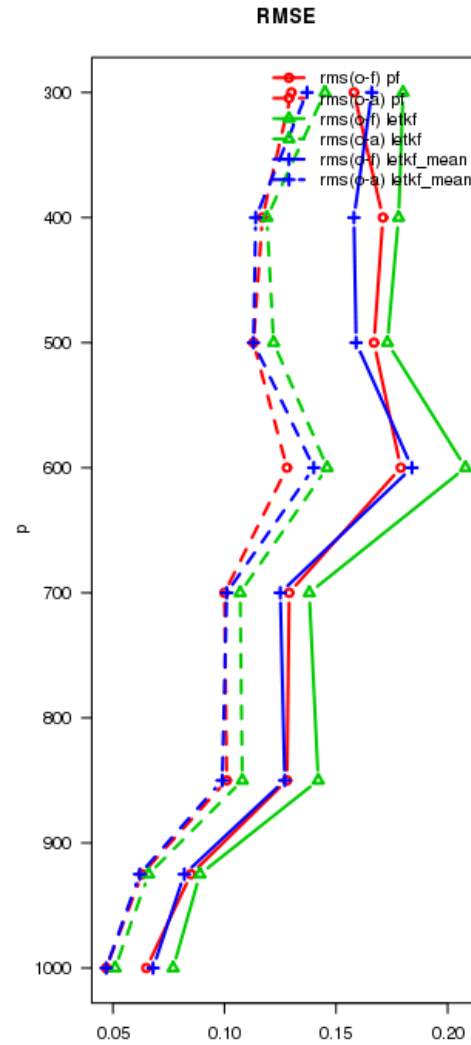
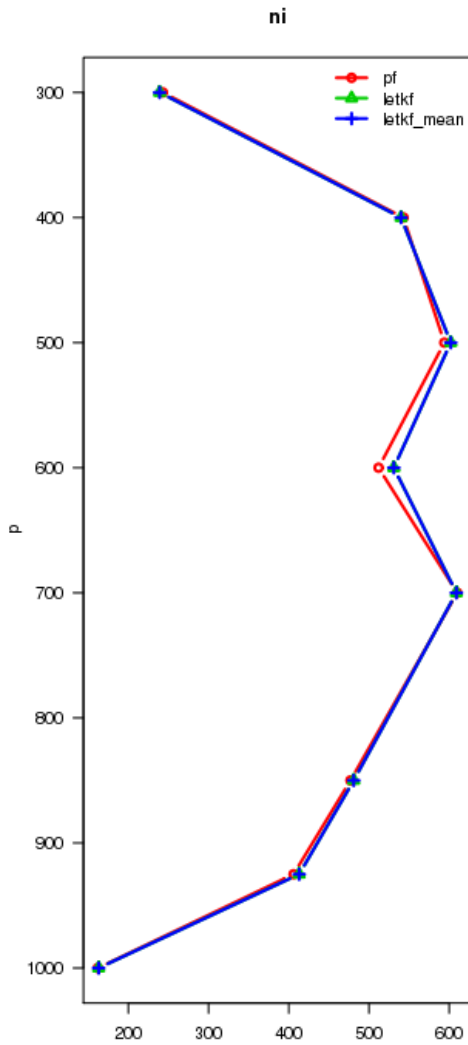


Humidity statistics for AIREP TEMP PILOT
experiments: pf, letkf, letkf_mean
startdate: 20160526130000 enddate: 20160527130000



LMCPF for COSMO

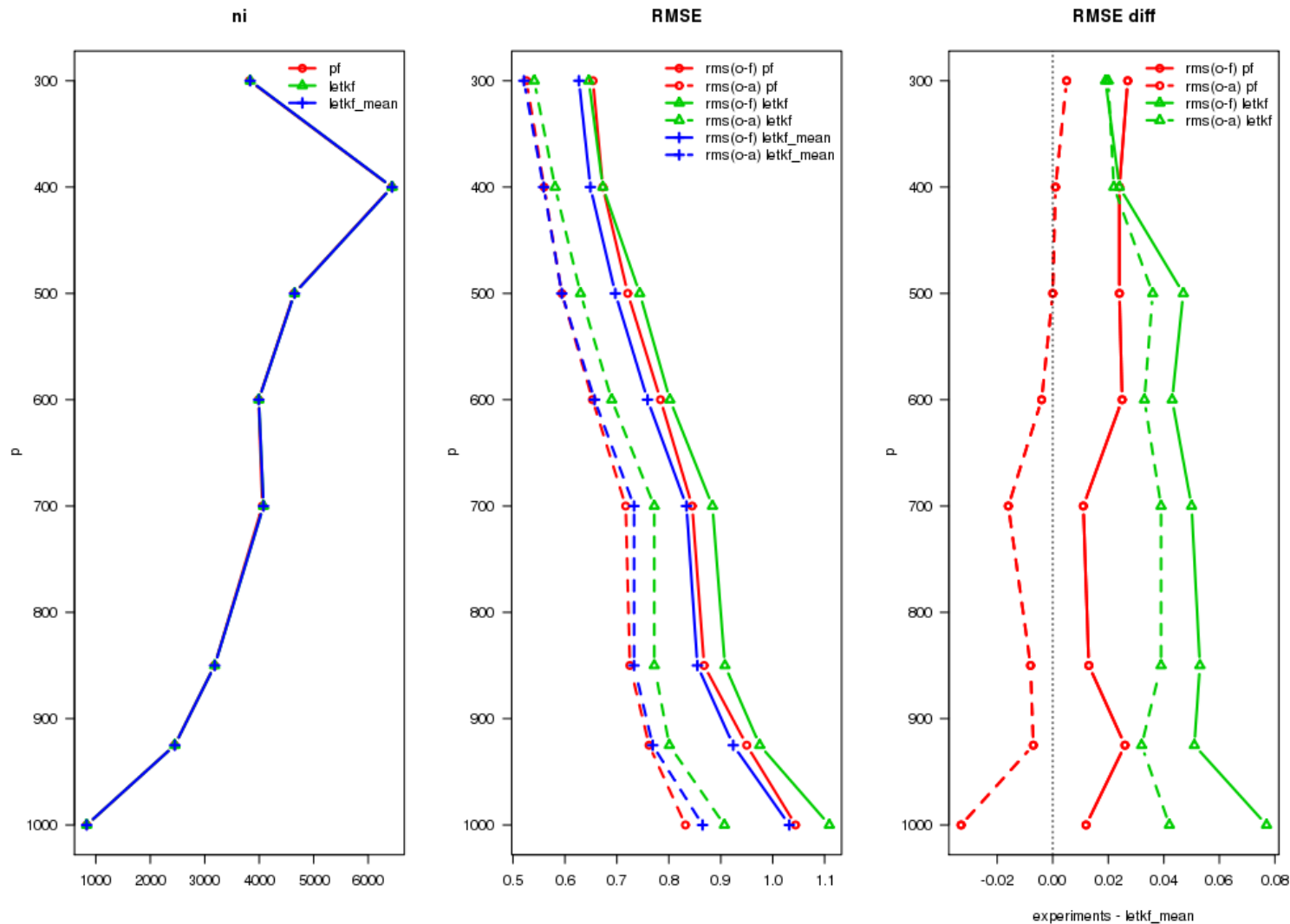
Humidity statistics for AIREP TEMP PILOT
experiments: pf, letkf, letkf_mean
startdate: 20160526130000 enddate: 20160527130000



LMCPF for COSMO

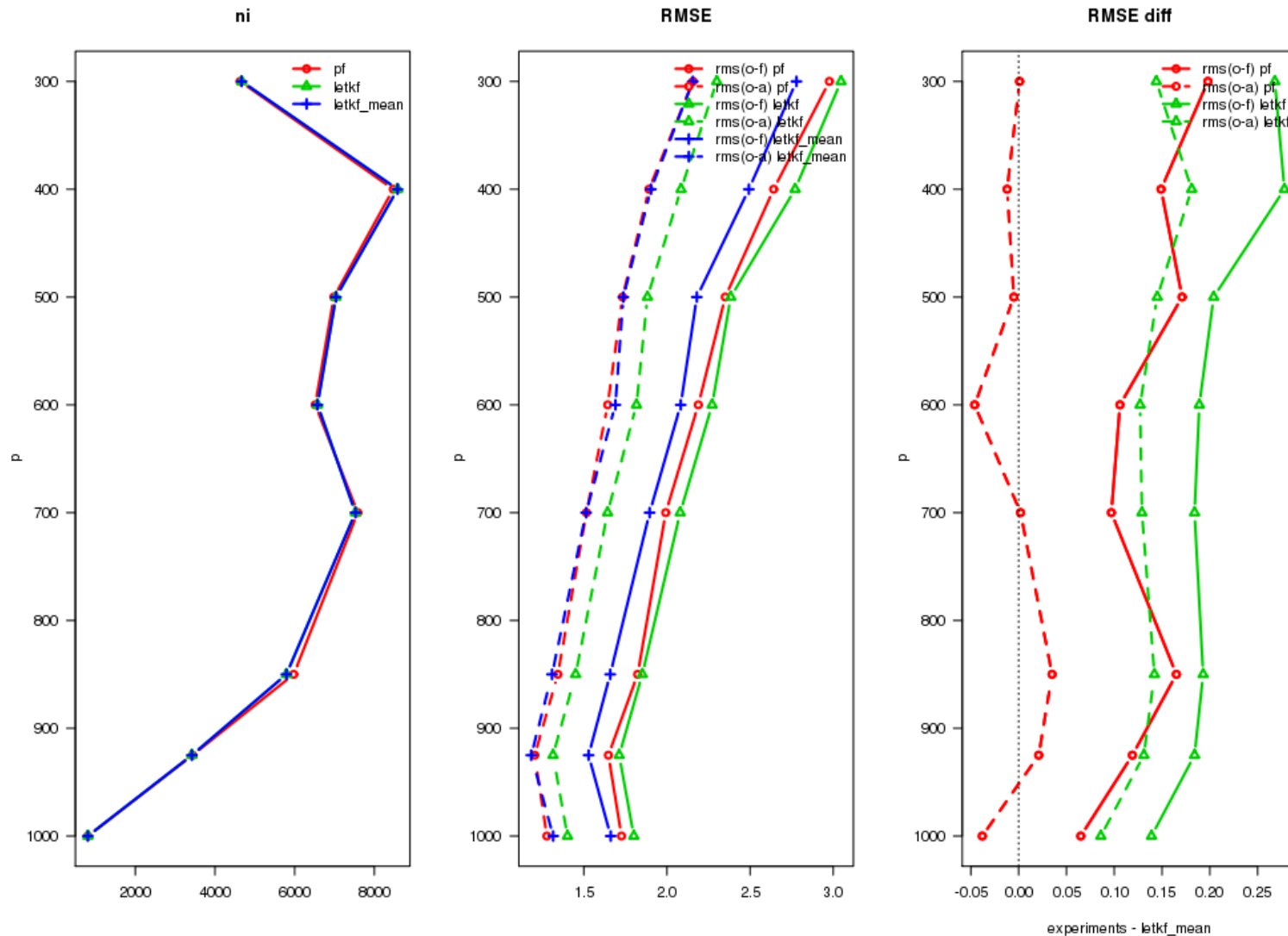
Temperature statistics for AIREP TEMP PILOT

experiments: pf, letkf, letkf_mean
startdate: 20160526130000 enddate: 20160527130000



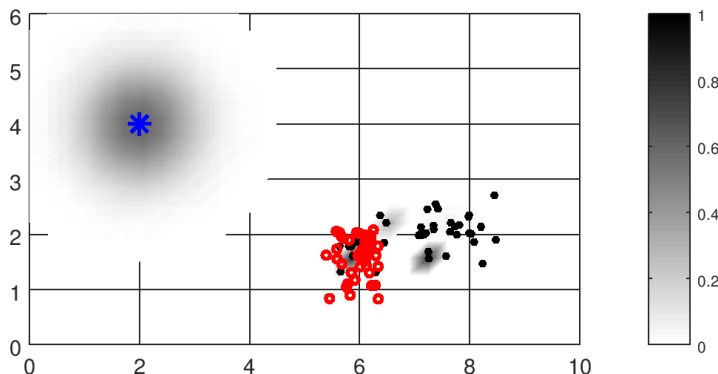
LMCPF for COSMO

Wind statistics for AIREP TEMP PILOT
experiments: pf, letkf, letkf_mean
startdate: 20160526130000 enddate: 20160527130000

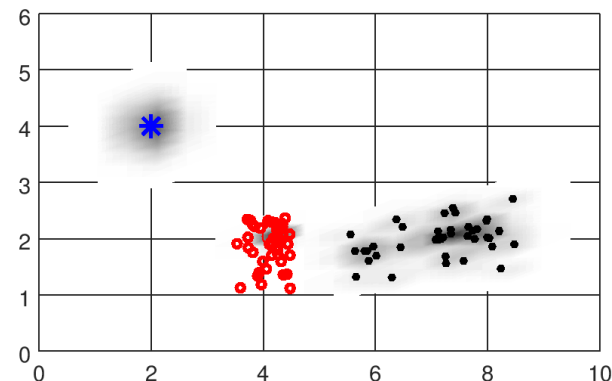


Summary LAPF and LMCPF

LAPF



LMCPF



- LAPF and LMCPF are implemented in an **operational NWP system**: **Globally + mesoscale + convective scale (KENDA)**
- Both Particle Filters are able to provide **reasonable atmospheric analysis** in a large-scale (high-dimensional) environment and are running stably over a period of one month
- The LMCPF outperforms the LAPF, both Particle Filters are not far behind the operational LETKF, **LMCPF starting to be comparable**

Both Particle Filters are showing promising results; further tuning and development is in progress.

Many Thanks!



Inverse Modeling

An introduction to the theory and methods of inverse problems and data assimilation

Gen Nakamura
Roland Potthast

